Spire Fall 2021 Ultimate Notebook

Contributions

Ben:

- Exploratory data analysis of the CYGNSS Full DDM Lv1 data
- Prepare data for interpolation/finish the collocation of interpolated data
- Prepare the modeling dataset with all relevant DDM calibration variables
- Linear regression modeling

David:

- Exploratory data analysis of the ECMWF background wind speed and significant wave height (SWH)
- Interpolating background wind speed and SWH for specular points
- · Machine learning modeling

Required Files:

- The only Python package files which needed to be installed for the project (weren't already recognized by Colab) were in the Haversine package
- Input files include several .nc (NetCDF) or nc4(netCDF4) files, a few .pkl files, and one .png file (involved in our Template Matching process)
- The team produced a database of collocated data, in 45 netCDF files, a shortcut to which will be included in the Spire project folder

Abstract

Ever since the launching of the CYGNSS satellite system in 2016 by NASA, many data analysts have experimented with models that use its Delay Doppler Map (DDM) information to predict weather phenomena. Together, team members of the Fall 2021 Spire project team have analyzed DDM datasets from CYGNSS, interpolated wind speed and wave height data from the European Centre for Medium-Range Weather Forecasts (ECMWF), collocating those with corresponding CYGNSS data. They have performed both linear and machine learning modeling on the collocated sets. The present analysis offers preparation, diagnostics and conclusions for both models, including functions for interpolation/collocation of ECMWF and CYGNSS data along the way.

Introduction

• What is the motive and history (before Math Clinic) of the project?

NASA'S CYGNSS (CYCLONE GLOBAL NAVIGATION SATELLITE SYSTEM) program was launched in 2016. It consists of eight satellites devoted to the gathering of weather data over the world's oceans, with the aim of providing meteorologists and data scientists input for models that will better predict the emergence of hurricanes and tropical cyclones. The satellites operate on a bistatic spacial priciple, being grouped into pairs, each pair consisting of a transmitter satellite and a receiver satellite. Data is collected through a process of reflectometry: the transmitter sends radio waves to the surface of the ocean at a certain doppler frequency. The amount of time it takes for the signal to reflect on the ocean surface at a 'specular point' and return to the receiver, combined with the shift in doppler frequency, both of which are functions of wind speed and wave height over the surface, are variables used to compute 'raw counts' power values. These values, in turn, are used to color Delay Doppler Maps (DDMs), 4 per timestamp/CYGNSS sample; all these maps are stored in netCDF files by sampling date, in a massive CYGNSS database, most of which is freely available for download through NASA's OpenDap system. Further information about the history and methods of the CYGNSS program can be found here: Source

The major motivation for the project is a desire, on the part of project sponsor Spire Global, to have wind speed/wave height data from the European Center for Medium-range Forecasting (ECMWF) background grids collocated with the CYGNSS Full DDM data offered by NASA for dates between March 1 and September 1, 2021, perhaps for use in future modeling. As an additional point of academic curiosity, the team was motivated to perform some elementary modeling with an eye toward predicting wind speed and wave height from calibrations of the DDM data.

• What (ideally) will be its impact on the sponsor or other stakeholders?

Ideally, the impact of the project on the sponsor will be to provide their data analysts with a pool of collocated data for any use in future modeling involving CYGNSS DDMs for 2021. Perhaps the collocated database will even prove useful in helping future Math Clinic groups that are working with the CYGNSS data in their efforts to contribute to forward model development, as is described in Huang, 2020.

• What is the general technical approach to solve the sponsor's problem?

The process of analyzing and preparing data for any eventual modeling began with independently exploring CYGNSS Full DDM datasets. Then, the team moved into a phase of interpolating ECMWF wind data/wave data (provided by Spire) by the location of specular points of CYGNSS sets against the background wind data/wave data grid. The data was interpolated using inverse distance weighting interpolation and the distance between points was found using the haversine forumla. The collocation process was then completed for 45 CYGNSS files between March 1 and September 1, 2021, in an attempt to process as much data in that time frame as possible, per the request of the project sponsors. Finally, Linear Regression and Machine Learning Modeling was conducted independently by separate team

members, with separate models predicting wind speed/wave height based on input from DDM calibrations. This involved considerable research and diagnostic processing. At the end of the modeling process, the model results were analyzed and their limitations detailed.

Methods

Data

• Describe the data or computational space, their amount (in bytes, number of files etc.), type (categorical, numerical etc.) and physical units

Project data came from NASA's CYGNSS Delay Doppler Map data and from ECMWF background wind/wave grid data, provided by the sponsor. Altogether, the team processed tens of gigabytes of that data in the collocation phase, during which team members built a collocated database of 45 files. The team then went on to perform modeling on a dataset built from 5 of those collocated files. Physical units included meters/sec for wind speed values, meters for wave height, datetime64 datatype units down the half-second for CYGNSS UTC timestamps, and delay times and doppler frequencies for power values that color delay doppler maps.

The teams extensive use of CYGNSS DDMs merits a more detailed definition of Delay Doppler Maps. According to a paper by University of Michigan and Soutwest Research Institute scholars Randy Rose, Scott Gleason and Chris Ruf, "A perfectly smooth surface reflects a specular point while a rough surface scatters it across a distributed "glistening zone". The Delay Doppler Map (DDM) created by the GNSS-R instrument is an image of that scattering cross-section in the time and frequency domains across the glistening zone" (Gleason, et. al, 2014).

**paper can be found at: Link

More specifically, it appears that the receiving device measures the time delay and doppler frequency of the reflected signal and cross correlates them with a "local copy" of original values for those variables from the transmitter. That correlation function is given in Huang, et, al, 2020 as:

$$Y_k(\tau, f) = \frac{1}{T_i} \int_{t_k - T_i}^{t_k} u_r(t) a(t + \tau) e^{2pij(f_0 + f)t} dt$$

where τ is the time delay, f is the frequency measure, t_k is the time for the "complex correlation result", a(t) is a function of the PRN code (also given in this dataset) and T_i is the integration time (in CYGNSS' case, $T_i = 1$ ms).

The N=1000 sequential results are then "incoherently averaged":

$$Z(\tau, f) = \frac{1}{N} \sum_{k=1}^{N} |Y_k(\tau, f)|^2$$

The power values stored in "raw_counts" in this dataset are just a linear combination of $Z(\tau,f)$

All this can be found in the Huang (2020) paper: "A Forward Model for Data Assimilation of GNSS Ocean Reflectometry Delay-Doppler Maps" Link

According to Gleason, et. al, "The DDM is an information-rich data set of surface state statistics. When this measurement is obtained from the ocean's surface, the data is intimately related to the surface wind vector and providing a direct measurement of the wave height statistics." The authors continue: "In the case of ocean surface GNSS scatterometry, estimation of the ocean surface roughness and near-surface wind speed is possible from two different properties of the DDM: The maximum scattering cross-section (the dark red region...) and the shape of the scattering arc [yellow/red regions]..." (Gleason, 2014).

Important DDM calibrations for modeling that were calculated by the team were three in number. The first was DDM Average, a simple average of 'raw counts' power values that color DDM plots in a 10 X 5 area around the specular point bin of each DDM. The second was RMS ratio, the highest power value in a given DDM divided by the root mean square of the rest of the power values (a statistic recommended for inclusion by the instructor). A third was Maximum Template Matching Coefficient, the highest pixel correlation coefficient in a template matching of each DDM against an 'ideal' template. The template matching method chosen was the Normalized Correlation Coefficient Method in the Open CV package in Python. This method works, according to Open CV documentation, by sliding across T, the template image matrix of pixels, and across I, the source/testing image matrix of pixels, and compares pixels, calculating a correlation metric, in our case: $R(x,y) = \frac{\sum_{x',y'} (T'(x',y') * I'(x+x',y+y'))}{\sqrt{\sum_{x',y'} T'(x',y')^2 * \sum_{x',y'} I'(x+x',y+y')^2}}$ Link

Another important pair of variables in our dataset were Normalized Bistatic Radar Cross Section (NBRCS) and Leading Edge Slope (LES), NASA's own, more sophisticated physical calibrations of the DDMs in the CYGNSS datasets. These were pulled directly from NASA's publically available 'ALL DDM' dataset, and then they were worked back into our modeling dataset.

The team had limited time to understand NBRCS and LES in all their wave-physical details, but outside research led team members to conclude they might be useful values in modeling. Methods of calibration of NBRCS and LES are described in the following article in the journal *Remote Sensing*: <u>Link</u>

The ECMWF dataset includes information about wind speed and wave heights for latitude and longitude pairs measured in 1/8 degree incriments. The variables of interest from the ECMWF data sets were 'U10m' which is a measurement of the 10 meter zonal wind (m/s), 'V10m' which is a measurement of the 10 meter meridonal wind (m/s), and 'SWH' which is a measurement of significant wave height.

In most of the world the standard is to measure wind speeds at a height of 10 meters above ground level. This ensures the measurement is not affected by surrounding vegitation. The United States measures at 20 feet above ground level. The wind is measured at that height to get a measurement unobstructed by other objects on the ground. Link

Significant wave height is a measurement devised by Walter Munk during World War II that measures the average wave height from through to crest of the highest 1/3 of waves. Significant wave height is used to estimate many aspects of a wave. The top 10% of waves are roughly 1.3 times SWH and the maximum wave height one would expect to see is roughly double the SWH. <u>Link</u>

Interpolation of Background Data

Interpolation is a technique used to estimate a property of an area based on known values of that property in surrounding areas. This uses the assumption that the properties behave more like areas nearby than areas far away. We often do not have the technology to collect data for a continuous region so things like weather stations will collect data and that data is interpolated to estimate the weather in areas around the weather station. The best method for interpolating geospatial data depends on a what weather property you are trying to interpolate. Based on the research the team found, there are conflicting opinions as to what is the best method for interpolating wind speed data. The most common method suggested was some form of inverse distance weighted interpolation. Acording to esri, a top GIS company, the convention for using inverse weighted distance interpolation with geospatial data is to use inverse weighted distance squared interpolation. (Link) The formula for inverse weighted distance squared interpolation is given by:

$$Z(x) = \frac{\sum_{i=1}^{n} (\frac{z_i}{d_i^2})}{\sum_{i=1}^{n} (\frac{1}{d_i^2})}$$

where z(x) is the interpolated value at point x, z_i is a known value of z and d_i^2 is the squared distance from point i to point x.

Finding the distance between two points on the globe is done using the <u>haversine</u> formula. This uses triginometry to find the distance between two points on a sphere and is commonly used to calculate the distance between two points on the globe. The haversine formula for is iven by:

$$d = 2rsin^{-1}(\sqrt{sin^2(\frac{\phi_2-\phi_1}{2}) + cos(\phi_1)cos(\phi_2)sin^2(\frac{\lambda_2-\lambda_1}{2})})$$

where d is the distance between two points, r is the radius of the sphere, in this case that is the radius of the Earth, ϕ_1 and ϕ_2 are the latitude of the two points and λ_1 and λ_2 are the longitude of the two points.

What mathematical, statistical, or physical models are being used?

The two main types of models being used are linear regression models (ordinary least squares) and machine learning models.

Modeling

Linear Regression:

Diagnostic work for the linear regression models involved the use of variance inflation factor analysis for multicolinearity assessment. Best subsets variable selection was performed, emphasizing the maximization of Residual Sum of Squares values from various models fit. Error assumptions were checked with normality plots to determine distribution of residuals, and with Durban-Watson test statistics to assess autocorrelation of consecutive errors in the model.

Outliers were examined using a Bonferroni Test. According to *Linear Models with R*, by Julian Faraway, we looked for observations whose Bonferonni corrected p-value, here called 'bonf(p)', is less than the studentized residual for the observation, where the studentized residual is equal to

$$t_i = r_i (\frac{n-p-1}{n-p-r_i^2})^{1/2}$$
,

and where r_i is the residual for that observation, n is the number of observations, and p is the number of regressors in the model. In the Bonferonni test, the Bonferonni adjusted p-value is equal to $\frac{\alpha}{n}$, where $\alpha=.05$.

Influential values were analyzed using Cook's Distance statistics, plotted against 'instances' (individual observations ordered by sample/time). The formula for Cook's Distance is given in Faraway's text as: $D_i = \frac{1}{p} r_i^2 \frac{h_i}{1-h_i}$, where p is the number of regressors in the model, r_i^2 is the residual effect of observation i squared, and $\frac{h_i}{1-h_i}$ is the 'leverage term' for the observation. This can be found in Faraway, Julian. Linear Models with R: 2nd ed., UK: CRC Press, 2015, pages 90-91.

Model structure was evaluated using fitted values vs. residuals plots and partial regression plots, to determine which DDM calibrations had the most significance in the model fit for predicting wind speed and wave height. Lastly, models were fit and their estimated coefficients interpreted in the context of the problem of predicting wind speed/wave height.

Machine Learning:

Machine learning is one of the most exciting and rapidly growing fields in the world. There is more depth to this than can be grasped in such a small amount of time. The general goal of a machine learning algorithm is minimizing a cost function. That cost function is a function that determines how close the model was to correctly predict test observations. The model is able to learn how to make its guesses by inputing training data. In general, the more training data available, the more accurate the models' predictions will be.

Since the Spire team aims to use the collocated data base to help train a machine learning model, and there are so many basic machine learning functions that are ready to use with minimal understanding, the team decided to attempt to develope a basic classification machine learning model.

A classification machine learning model takes in observations that train the model to recognize a class that observation belongs in. After the model is trained, it tries to predict what class an observation should be in. A common method for optimizing a classification model is to use gradient decent. Gradient decent is a great tool for minimizing a cost function. Stochastic gradient descent uses a single sample to

compute the gradient. Sci-learn kit offers a Stochastic gradient classification alorithm called SGD-Classifier. This is the alorithm the team will use to create their models.

Code for Notebook setup

Loading libraries

```
1 # Install necessary packages
 2 !pip install haversine
    Collecting haversine
      Downloading haversine-2.5.1-py2.py3-none-any.whl (6.1 kB)
    Installing collected packages: haversine
    Successfully installed haversine-2.5.1
 1 ##import necessary packages
 2 import xarray as xr
 3 import netCDF4 as nc
 4 import pandas as pd
 5 import numpy as np
 6 import matplotlib
 7 import matplotlib.pyplot as plt
 8 import pylab as py
 9 import itertools
10 import seaborn as sns
11 import os
12 from matplotlib import ticker
13 from matplotlib import colors
14 import matplotlib.patches as mpatches
15 import math
16 from datetime import *
17 from datetime import datetime
18 from scipy.stats import halfnorm
19 from scipy.special import cbrt
20 from patsy import dmatrices
21 import statsmodels.api as sm
22 from statsmodels.stats.outliers influence import variance inflation factor
23 from yellowbrick.base import Visualizer
24 from statsmodels.stats.outliers_influence import OLSInfluence as influence
25 from sklearn.linear_model import LinearRegression
26 import cv2
27 from google.colab.patches import cv2_imshow
28 from sklearn.model_selection import train_test_split
29 from sklearn.metrics import mean_squared_error
30 import pickle
31 from os.path import exists
32 from haversine import haversine, Unit
```

```
33 from tqdm import tqdm
34 from sklearn.linear_model import SGDClassifier
35 from sklearn.datasets import load iris
36 from sklearn.datasets import make classification
37 from sklearn.model selection import train test split
38 from sklearn.metrics import confusion matrix
39 from sklearn.metrics import classification report
40 from sklearn.preprocessing import scale
41 from statsmodels.graphics.mosaicplot import mosaic
42 from matplotlib.patches import Patch
43 import itertools
44 from collections import deque
45 %matplotlib inline
46 %precision 3
    /usr/local/lib/python3.7/dist-packages/statsmodels/tools/_testing.py:19: FutureWarnin
      import pandas.util.testing as tm
    '%.3f'
```

Drive mount/File path variables

```
1 #Allow user to mount to user google drive
2 from google.colab import drive
3 drive.mount("/content/drive")

    Mounted at /content/drive

1 #Names of respective shortcuts in the Spire folder to team member Google drive file fol
2 David_path = 'For Davids Notebook v2/'
3 Ben_path = 'Files needed to run Spire project notebook/'

1 #Define variables to be used later in i/o pathways
2 cwd = os.getcwd()  # Assumes no cd commands were executed
3 pathProfessor = 'Colab Notebooks/Math Clinic/2021fa/Spire/'
```

Functions for Notebook Execution

```
##FUNCTIONS FOR EXPLORATORY ANALYSIS PHASE:

## define a function that plots the CYGNSS latitudes against CYGNSS longitudes and col
def location_time_plot(ultimate_set,start_sample_index,end_sample_index):
    selection = ultimate_set.isel(sample=slice(start_sample_index,end_sample_index))
    times_array = selection['ddm_timestamp_utc'].values
    times_array = times_array.flatten()
    times_array = convert_to_minute(times_array)
    times_array = np.repeat(times_array, 4)
```

```
10
     lons = selection['sp lon'].values
11
     lons = lons.flatten()
     lats = selection['sp lat'].values
12
13
     lats = lats.flatten()
     cmap = matplotlib.cm.get cmap("viridis", 2)
14
     ax1 = plt.scatter(x=lons, y = lats, s = 3, c = times array, cmap = cmap)
15
16
     cb = plt.colorbar(ax1)
     tick locator = ticker.MaxNLocator(nbins= 2)
17
     cb.locator = tick locator
18
19
     cb.update ticks()
     cb.ax.set yticklabels(['dropped string','1:05-1:15', '17:02-17:12'])
20
     plt.xlabel('Specular Point Longitude')
21
    plt.ylabel('Specular Point Latitutde')
22
     plt.title('Specular Point Latitude vs. Longitude, Colored by Minute since 00:00, 4-1
23
24
25 #define a function that will convert the times array in function 'location time plot'
26 def convert_to_minute(times_array):
27
     times array1 = [str(stamp) for stamp in times array]
28
     times array2 = [stamp[11:16] for stamp in times array1]
29
     times_array_nump = np.array(times_array2)
     times array nump = np.char.replace(times array nump, ':', '')
30
31
     hour array = [stamp[0:2] for stamp in times array nump]
32
     minute array = [stamp[2:4] for stamp in times array nump]
33
     hour integer = [int(stamp) for stamp in hour array]
34
     hour in min = [element * 60 for element in hour integer]
35
     minute integer = [int(stamp) for stamp in minute array]
36
     sum list = [a + b for a, b in zip(hour in min, minute integer)]
37
38
     return sum list
39
40 #define a function that will plot all four DDMs associated with a given CYGNSS sample
41 def ddm_plots(ourset, sample_select):
42
    fig, axes = plt.subplots(2,2)
43
     plt.suptitle(f'DDMs for Sample {sample select}', va = 'bottom')
     ourset.sel(sample = sample_select, ddm = 0)['raw_counts'].plot(ax = axes[0][0], add_
44
45
     ax = axes[0][0]
46
     ax.set title('Channel 0')
47
     ourset.sel(sample = sample select, ddm = 1)['raw counts'].plot(ax= axes[0][1], add 1
48
     ax = axes [0][1]
49
     ax.set title('Channel 1')
     ourset.sel(sample = sample select, ddm = 2)['raw counts'].plot(ax= axes[1][0], add 1
50
51
     ax = axes[1][0]
52
     ax.set title('Channel 2')
     ourset.sel(sample = sample select, ddm = 3)['raw counts'].plot(ax= axes[1][1], add 1
53
54
     ax = axes[1][1]
55
     ax.set title('Channel 3')
56
     fig.tight_layout()
57
58 def latlon index(ds, x, y, tol = 1/8):
59
60
    Function that returns index values of nearest grid point to longitude
61
     and latitude coordinates
62
     ds = dataset
     x = latitude coordinate (W < 0 <= E) by convention
63
```

```
y = longitude coordinate (S < 0 <= N) by convention
 64
 65
      tol = The tolerance of the measuring device (distance between grid points)
 66
       1 1 1
 67
      idx x = (ds.lat > x - tol/2) & (ds.lat <= x + tol/2)
 68
 69
      idx x = np.where(idx x)[0]
 70
      idx x = idx x[0]
 71
      idx y = (ds.lon > y - tol/2) & (ds.lon <= y + tol/2)
 72
 73
      idx y = np.where(idx y)[0]
 74
      idx y = idx y[0]
 75
 76
      return(idx x, idx y)
 77
 78 def vizualize region(ds, x, y, tol = 1/8, grid size = 24, overlay=True, alpha = 1):
 79
 80
      Function that generates some basic plots to observe an area of interest
 81
      ds = dataset
      x = latitude coordinate (E > 0 & W < 0) by convention
 82
 83
      y = longitude coordinate (N > 0 & S < 0) by convention
      tol = The tolerance of the measuring device (distance between grid points)
 84
 85
      overlay = if True, Wind speed vector map and SWH contour plot will be overlain on ea
 86
      alpha = transparancy of SWH contour map if overlay is set to True
 87
 88
      Plot 1: v10m vs u10m
      Plot 2: lon vs lat
 89
      Plot 3: swh boxplot
 90
 91
      Plot 4: spd histogram
 92
      Plot 5: 10m Wind Speed vector field
 93
      Plot 6: Countour plot of SWH
 94
      1 1 1
 95
 96
      # returns latitude and longitude index values for the grid point that is
 97
      # closest to the given coordinates
      lat_idx = (ds.lat > x - tol/2) & (ds.lat <= x + tol/2)
 98
 99
      lat idx = np.where(lat idx)[0]
      lat idx = lat idx[0]
100
      lon idx = (ds.lon > y - tol/2) & (ds.lon <= y + tol/2)
101
102
      lon idx = np.where(lon idx)[0]
103
      lon idx = lon idx[0]
104
      # Creates varaibles for plotting
105
106
      x min = lat idx - int(grid size/2)
      x max = lat idx + int(grid size/2)
107
108
      y min = lon idx - int(grid size/2)
      y max = lon idx + int(grid size/2)
109
110
      u = ds['U10m'][x_min:x_max, y_min:y_max]
      v = ds['V10m'][x_min:x_max, y_min:y_max]
111
112
      lat = ds['lat'][x min:x max]
      lon = ds['lon'][y_min:y_max]
113
114
      swh = ds['SWH'][x min:x max, y min:y max]
115
      spd = ds['SPD'][x_min:x_max, y_min:y_max]
116
117
      # Variables for legend of Plot 2
```

```
grid leg = mpatches.Patch(color = 'blue', label = 'Grid Point')
118
119
      spec_leg = mpatches.Patch(color = 'red', label = 'Input coordinates')
120
121
      fig = plt.figure()
122
      fig.set size inches(20, 15)
123
124
      ax1 = plt.subplot2grid((3, 4), (0, 0))
125
      ax2 = plt.subplot2grid((3, 4), (0, 1))
      ax3 = plt.subplot2grid((3, 4), (0, 2))
126
127
      ax4 = plt.subplot2grid((3, 4), (0, 3))
      X, Y = np.meshgrid(lon, lat)
128
129
      # Plot 1
130
      ax1.scatter(u, v, s = 2)
131
132
      corr = round(float(xr.corr(u, v)), 2)
      ax1.set title(f'v10m vs u10m correlation = {corr}')
133
134
      ax1.set xlabel('u10m')
135
      ax1.set ylabel('v10m')
136
137
      # Plot 2
      ax2.scatter(X, Y, s = 3)
138
139
      ax2.plot(y, x, 'ro', markersize = 5)
140
      idx x, idx y = latlon index(ds, x, y)
      dist = round(haversine((x, y), (ds['lat'][idx_x], ds['lon'][idx_y])), 2)
141
142
      ax2.set title(f'Distance to nearest gridpoint: {dist} (km)')
      ax2.set xlabel('Longitude')
143
144
      ax2.set ylabel('Latitude')
      ax2.legend(handles=[spec leg, grid leg], frameon=True)
145
146
147
      # Plot 3
      swh df = swh.to dataframe()
148
149
      ax3.boxplot(swh df['SWH'])
150
      mean swh = round(np.mean(swh df['SWH']), 2)
151
      std swh = round(np.std(swh df['SWH']), 2)
      ax3.set_title(f'Avg SWH = {mean_swh:.2e} (m), std = {std swh}')
152
153
      ax3.set ylabel('SWH (m)')
154
155
      # Plot 4
      spd df = spd.to dataframe()
156
157
      ax4.hist(spd df['SPD'], bins=25)
      mean spd = round(np.mean(spd df['SPD']), 2)
158
159
      std spd = round(np.std(spd df['SPD']), 2)
160
      ax4.set xlabel('Windspeed (m/s)')
      ax4.set title(f'Avg windspeed = {mean spd} (m/s), std = {std spd}')
161
162
163
      if overlay == True:
        ax5 = plt.subplot2grid((3, 4), (1, 0), colspan=3, rowspan=2)
164
165
        # Plot 5
166
        plot1 = ax5.contourf(X, Y, swh, cmap='ocean', alpha=alpha)
167
        plot2 = ax5.quiver(X, Y, u, v, spd, cmap='jet')
168
        ax5.set title('Wind speed vector map and SWH contour map')
169
170
        ax5.set xlabel('Longitude')
        ax5.set ylabel('Latitude')
171
```

```
plt.colorbar(plot1, ax=ax5, label='SWH')
172
173
         plt.colorbar(plot2, ax=ax5, label='Wind Speed')
174
      else:
175
         ax5 = plt.subplot2grid((3, 4), (1, 0), colspan=2, rowspan=2)
         ax6 = plt.subplot2grid((3, 4), (1, 2), colspan=2, rowspan=2)
176
177
178
         # Plot 5
         plot5 = ax5.guiver(X, Y, u, v, spd, cmap='jet')
179
         fig.colorbar(plot5, ax=ax5, shrink=0.8)
180
181
         ax5.set title('10m Wind Speed Vector Field')
182
         ax5.set xlabel('Longitude')
183
         ax5.set ylabel('Latitude')
184
         ax5.set aspect('equal')
185
186
         # Plot 6
187
188
         plot6 = ax6.contourf(X, Y, swh, cmap='ocean')
         fig.colorbar(plot6, ax=ax6, shrink=0.8)
189
190
191
         ax6.set title('Significant Wave Height Contour')
         ax6.set xlabel('Longitude')
192
193
         ax6.set ylabel('Latitude')
194
         ax6.set aspect('equal')
195
196
      plt.subplots adjust(wspace = 0.3, hspace = 0.3)
197
      plt.show()
198
199 def latlon index(ds, x, y, tol = 1/8):
200
201
      Function that returns index values of nearest grid point to given longitude
      and latitude coordinates
202
203
      ds = dataset
204
      x = latitude coordinate (W < 0 <= E) by convention
      y = longitude coordinate (S < 0 <= N) by convention
205
206
      tol = The tolerance of the measuring device (distance between grid points)
207
       1 1 1
208
209
      idx x = (ds.lat > x - tol/2) & (ds.lat <= x + tol/2)
210
      idx x = np.where(idx x)[0]
211
      idx x = idx x[0]
212
      idx y = (ds.lon > y - tol/2) & (ds.lon <= y + tol/2)
213
214
      idx y = np.where(idx y)[0]
215
      idx y = idx y[0]
216
217
      return(idx x, idx y)
218
219 ##FUNCTIONS FOR INTERPOLATION PHASE:
220
221
    def spec_values_nearest(ds, sp_lat, sp_lon, tol = 1/8):
222
      Input an xarray dataset, latitude coordinate, longitdue coordinate, and tollerance
223
224
      Output U10m, V10m, and SWH of the grid point that is closest to the input coordinate
225
      ds = xarray dataset
```

```
sp lat = latitude coordinate of specular point
226
227
      sp_lon = longitude coordinate of specular point
      tol = tollerance of measuring device (distance between grid points)
228
229
       1 1 1
230
231
      #Finding index values of the nearest gridpoint
      idx x, idx y = latlon index(ds, sp lat, sp lon)
232
233
234
      # Assigning values of interest to variables
235
      U10m = float(ds['U10m'][idx x, idx y])
      V10m = float(ds['V10m'][idx x, idx y])
236
      SWH = float(ds['SWH'][idx x, idx y])
237
238
239
      return(U10m, V10m, SWH)
240
241 def clean CYGNSS(ds):
242
      In rare cases, CYGNSS data is not input correctly. One case was found where a
243
244
      specular point had a longitude greater than 180.
      This function checks for out of range longitude coordinates and corrects them.
245
        *Longitude of 190 is the same as longitude of -170.
246
247
248
      Currently this function does not check for or correct out of range latitude coordina
      This is because a latitude value over 90 or under -90 has not been found in a CYGNSS
249
250
      data set and does not make sense physically as that point does not exist on the glob
      If such a case occurs, this function will be adjusted to include that functionality.
251
252
253
254
      n = len(ds)
255
      # case: secular point longitude is greater than 180
256
257
      for i in range(len(ds)):
258
        if ds.at[i, 'sp lon'] > 180:
259
           ds.at[i, 'sp lon'] = -(360 - ds.at[i, 'sp lon'])
260
261
      # case: specular point longitude is less than 180
262
      for i in range(len(ds)):
        if ds.at[i, 'sp lon'] < -180:
263
           ds.at[i, 'sp_lon'] = (360 + ds.at[i, 'sp_lon'])
264
265
266
      return(ds)
267
   def idw interpolate(paired list, power = 2, default = np.nan):
268
      1 1 1
269
270
      Input:
271
      list in form [(d1, z1), \ldots, (dn, zn)] where:
272
        d1 = distance from first neighbor
        z1 = known value from first neighbor
273
274
        dn = distance from nth neighbor
275
        zn = distance from nth neighbor
      power = power for interpolation
276
277
      default = value used for missing or NA data
278
279
      Output: inverse distance weighting interpolated value
```

```
1 1 1
280
281
       num = [x[1]/(x[0]**power)] for x in paired list] # Generates list with each grid poin
282
       denom = [1/(x[0])**power) for x in paired list] # Generates list with each grid point
283
284
285
286
       if sum(denom) != 0:
        idw value = float(sum(num)/sum(denom))
287
288
       else:
289
        idw value = default
290
291
       return(idw value)
292
293 def interpolate point(ds, sp lat, sp lon, tol = 1/8, swh tol = 10000, power = 2, defau
294
       Input an xarray dataset, lattitude coordinate, longitude coordinate, and optional to
295
296
      Calculate disctance between the specular point and the four surrounding grid points
      haversine formula.
297
298
       Calculates wU10m, wV10m, wSWH using inverse distance squared weighted interpolation
299
      Returns weighted U10m (wU10m), weighted V10m (wV10m), and weighted SWH (wSWH)*
         *Some significant wave height values are entered as 1e20.
300
         To deal with this, grid points with SWH values > 10000 are not used for calculati
301
302
       1 1 1
303
304
       # Finds index values of 4 surrounding grid points
305
       idx x = np.where((ds.lat > sp lat - tol) & (ds.lat <= sp lat + tol))[0]
306
307
       idx y = np.where((ds.lon > sp lon - tol) & (ds.lon <= sp lon + tol))[0]
308
       # If the specular point lon is in the range (179.875, 180), the index value for -180
309
310
      if len(idx y) == 1:
311
        idx_y = np.append(idx_y, 0)
312
       # Finding inverse distance squared between spec point and 4 nearest grid points
313
       d1, d2, d3, d4 = (haversine((sp_lat, sp_lon), (ds['lat'][idx_x[0]], ds['lon'][idx_y[
314
315
316
       # Finding weighted values for U10m and V10m
       U10m list = [(d1, ds['U10m'][idx x[0], idx y[0]]), (d2, ds['U10m'][idx x[1], idx y[1])]
317
318
                    (d3, ds['U10m'][idx_x[0], idx_y[1]]), (d4, ds['U10m'][idx_x[1], idx_y[0
319
       V10m list = [(d1, ds['V10m'][idx x[0], idx y[0]]), (d2, ds['V10m'][idx x[1], idx y[1])]
320
                    (d3, ds['V10m'][idx_x[0], idx_y[1]]), (d4, ds['V10m'][idx_x[1], idx_y[0
       temp SWH = [(d1, ds['SWH'][idx x[0], idx y[0]]), (d2, ds['SWH'][idx x[1], idx y[1]])
321
                    (d3, ds['SWH'][idx x[0], idx y[1]]), (d4, ds['SWH'][idx x[1], idx y[0]]
322
323
       SWH list = [x \text{ for } x \text{ in temp SWH if } x[1] < \text{swh tol}] # Generates list with SWH values
324
325
326
      wU10m = idw interpolate(U10m list)
      wV10m = idw_interpolate(V10m_list)
327
328
      wSWH = idw interpolate(SWH list, default = default)
       U10m_neighbor, V10m_neighbor, SWH_neighbor = len(U10m_list), len(V10m_list), len(SWH
329
330
331
       return(wU10m, wV10m, wSWH, U10m_neighbor, V10m_neighbor, SWH_neighbor)
332
333 def interpolate date(mm, dd, power = 2, subset = False, len subset = 500, default = np
```

```
334
335
      mm is month (do not enter leading 0's e.g. January is 1 not 01)
      dd is the date (do not enter leading 0's e.g the first is 1 not 01)
336
337
      power = the power used for IDW interpolation
338
339
      This function takes in a date,
340
      Loads the appropriate datasets, and
      Generates a data frame with:
341
      index value
342
343
      specular point latitude
      specular point longitude
344
345
      weighted U10m (based on spec values() function)
      weighted V10m (based on spec values() function)
346
347
      weighted SWH (based on spec values() function)
348
      for all coordinates in the CYGNSS dataset
      Then it turns the data frame into a pickle and writes it to the drive
349
      output file: "/content/drive/MyDrive/Spire data/wValues/wValues 2021mmdd.pkl"
350
351
352
      Input files:
353
      location: "/content/drive/MyDrive/Spire data/For Notebook/file name"
      CYGNSS file in form: CYGNSS mmdd.pkl
354
355
        This file is a pickle formed from the xarray dataset for the input date
356
        The pickle file has the columns: 'Timestamp', 'specular point lat', 'specular poin
      Windspeed data file in form: 2021mmdd tod.nc
357
358
        * tod = 00, 06, 12, or 18 depending on time of day sampling was done
359
      subset and len subset variables allow for testing new functionality on small samples
360
      1 1 1
361
362
      # Variables for file path
363
      date = str(mm).zfill(2) + str(dd).zfill(2)
364
365
366
      # Loding CYGNSS file for input date and corrects coordinates that are out of bounds
367
      cyg file = f"{pathTeam}CYGNSS {date}.pkl"
      cyg = pd.read_pickle(cyg_file)
368
369
      cyg = clean CYGNSS(cyg)
370
371
      # Loading background windspeed data
      ds00 = f"{pathTeam}ecmwf.t00z.pgrb.0p125.f000 2021{date}00.nc"
372
373
      if exists(ds00):
        ds00 = xr.open dataset(ds00)
374
        print('ds00 was loaded')
375
376
      ds06 = f"{pathTeam}ecmwf.t06z.pgrb.0p125.f000 2021{date}06.nc"
      if exists(ds06):
377
        ds06 = xr.open dataset(ds06)
378
379
        print('ds06 was loaded')
      ds12 = f"{pathTeam}ecmwf.t12z.pgrb.0p125.f000_2021{date}12.nc"
380
      if exists(ds12):
381
382
        ds12 = xr.open dataset(ds12)
        print('ds12 was loaded')
383
384
      ds18 = f"{pathTeam}ecmwf.t18z.pgrb.0p125.f000_2021{date}18.nc"
385
      if exists(ds18):
386
        ds18 = xr.open dataset(ds18)
        print('ds18 was loaded')
387
```

```
388
      # Defining variables
389
      wcolumn names = ['lat', 'lon', 'wU10m', 'wV10m', 'wSWH']
390
391
      wValues = pd.DataFrame()
      neighbor col = ['U10m neighbor', 'V10m neighbor', 'SWH neighbor']
392
      neighbor count = pd.DataFrame()
393
394
      if subset == True:
395
396
        n = len subset
397
        date = date + str(' sample')
398
      else:
        n = len(cyg)
399
400
401
      # Loop that creates data frame with weighted values
402
      for i in tqdm(range(0,n)):
403
        x = float(cyg.iloc[i][1])
404
        y = float(cyg.iloc[i][2])
405
        # Using the hours in the Timestamp to determine what dataset to load
406
407
        ds tup = (ds00, ds06, ds12, ds18)
        ts = cyq.iloc[i][0].hour
408
409
        ts idx = int((ts - ts%6)/6)
410
        ds = ds tup[ts idx]
411
412
        # Calculating wU10m, wV10m, and wSWH
        wU10m, wV10m, wSWH, u10m neigh, v10m neigh, swh neigh = interpolate point(ds, x, y
413
414
415
        # Generating data frame
416
        wtemp df = pd.DataFrame([[x, y, wU10m, wV10m, wSWH]], columns = (wcolumn names))
417
        wValues = wValues.append(wtemp df)
        temp neighbor = pd.DataFrame([[u10m neigh, v10m neigh, swh neigh]], columns = (nei
418
419
        neighbor count = neighbor count.append(temp neighbor)
420
      filename = f"{pathTeam}wValues 2021{date}.pkl"
421
      wValues.to_pickle(filename)
422
423
      # Printing summary of data to check for errors
424
425
426
      print()
427
      print(wValues.describe())
428
      print()
      print('Neighbor count for U10m:')
429
430
      print(neighbor count['U10m neighbor'].value counts().sort index())
      print('Neighbor count for V10m:')
431
      print(neighbor count['V10m_neighbor'].value_counts().sort_index())
432
433
      print('Neighbor count for SWH:')
      print(neighbor count['SWH neighbor'].value counts().sort index())
434
435
436 def ecmwf check(mm, dd):
437
      mm = month (do not enter leading 0's e.g. January is 1 not 01)
438
      dd = date (do not enter leading 0's e.g the first is 1 not 01)
439
440
441
      Input a month and date
```

```
442
      Output a statement with a list of strings potentially containing: {'00', '06', '12',
443
      The strings represent the timeframe of ecmwf data needed to interpolate the data for
      the given date
444
445
      This function was run before 'idw interpolate' everytime to ensure the correct files
446
447
       1 1 1
448
      date = str(mm).zfill(2) + str(dd).zfill(2)
449
      ds = pd.read pickle(f"{pathTeam}CYGNSS {date}.pkl")
450
451
452
      ts list = []
      ts final = []
453
454
455
      for i in range(0, len(ds)):
456
       ts = ds.iloc[i][0].hour
457
        ts = ts - ts\%6
458
        ts = str(ts).zfill(2)
459
        ts list.append(ts)
460
461
      [ts_final.append(n) for n in ts_list if n not in ts_final]
462
463
      for j in range (0, len(ts final)):
464
        print(f'Need ECMWF file: ecmwf.t{ts final[j]}z.pgrb.0p125.f000 2021{date}{ts final
465
466 ##FUNCTIONS FOR COLLOCATION PHASE:
467
468 #define a function that will collect specular point latitudes/longitudes and timestamp
469 ##the function returns a pandas dataframe with the collected information
470 def collect latlons(cyg data set):
471
      time array = cyg data set['ddm timestamp utc'].values
472
      time array = time array.flatten()
473
      time array = np.repeat(time array, 4)
474
      lat array = cyg data set['sp lat'].values
475
      lat array = lat array.flatten()
      lon_array = cyg_data_set['sp_lon'].values
476
477
      lon array = lon array.flatten()
      temp_frame = pd.DataFrame({'timestamp':time_array, 'sp_lat':lat_array, 'sp_lon':lon
478
479
      return temp_frame
480
481 ##define function that integrates the collocated wind/wave data into the original xarr
482 def integrate sets(set to coll, coll set):
      coll set.reset index(drop= True, inplace = True)
483
484
      coll setA = coll set.drop(columns = ['lat', 'lon'])
485
486
      samp array = set to coll['sample'].values
487
      samp array.flatten()
488
      samp_array = np.repeat(samp_array, 4)
      ddm_array = set_to_coll['ddm'].values
489
490
      ddm array.flatten()
      ddm_array = np.concatenate((ddm_array, np.tile(ddm_array, 2415)))
491
492
      multi frame = pd.DataFrame({'sample':samp array, 'ddm':ddm array})
493
494
      mindx = pd.MultiIndex.from frame(multi frame)
495
      coll_setB = np.array(coll_setA)
```

```
multi frame coll = pd.DataFrame(coll setB, columns = ['wU10m', 'wV10m', 'wSWH'], ind
496
497
      collocated set = multi frame coll.to xarray()
      final coll set = xr.combine by coords([set to coll, collocated set])
498
499
500
      return final coll set
501
502 ##FUNCTIONS FOR DDM CALIBRATION PHASE:
503
504 ##define function that will calculate the ddm average for a 10x5 bin area around each
505 def DDM averages(cleaned set, sample first, delay start, delay end, doppler start, doppler
      isolate first = cleaned set.sel(sample = sample first, delay = range(delay start, del
506
      power first = isolate first['raw counts']
507
      DDM average first = power first.groupby('ddm').mean(dim=['delay','doppler'])
508
      average array = DDM average first
509
510
      for item in cleaned set['sample']:
511
        isolate = cleaned_set.sel(sample = item, delay = range(delay_start, delay_end), dopp
512
513
        power = isolate['raw counts']
        DDM average = power.groupby('ddm').mean(dim =['delay', 'doppler'])
514
515
        average_array = xr.concat([average_array, DDM_average], dim ='sample')
516
517
      average array = average array.drop duplicates(dim='sample')
518
      ds = average array.to dataset(name = 'ddm average')
      final set = xr.combine by coords([cleaned set, ds])
519
520
      return final set
521
522 def NBRCS LES vals(all data set, final set, start samp, stop samp, start of interval, end
      partition = final set.sel(sample = slice(start of interval, end of interval))
523
      partition2 = all data set.sel(sample = slice(start samp, stop samp))
524
525
      sampling array = partition['sample'].values
      sampling array = sampling array.flatten()
526
      partition2 = partition2.assign coords(sample = sampling array)
527
528
      nbrcs vals = partition2['ddm nbrcs']
      les vals = partition2['ddm les']
529
530
      ds1= nbrcs vals.to dataset(name='nbrcs')
      ds2= les vals.to dataset(name='les')
531
      result set = xr.combine by coords([partition, dsl], compat = 'override')
532
533
      result set = xr.combine by coords([result set, ds2], compat = 'override')
      return result set
534
535
536 ##define a function that finds the highest power value of DDM and divides it by the ro
    def root square ratio(complete set, sample, ddm channel):
537
538
      test ddm = complete set.isel(sample = sample, ddm = ddm channel)
      spec point select = test ddm.sel(delay = 64, doppler = 10)
539
      spec point power = spec point select['raw counts'].max().values
540
541
      test counts = test ddm['raw counts']
      test counts = test counts.values
542
543
      test counts = test counts.flatten()
      test counts = np.delete(test counts, np.argwhere(test counts == spec point power))
544
545
      squared = np.square(test counts)
      square sum = squared.sum()
546
      square sum divide = square sum/squared.size
547
548
      RMS = math.sgrt(square sum divide)
549
      RMS ratio = spec point power/RMS
```

```
550
      return RMS ratio
551
552 ##define a function that builds the RMS ratio index dataframe for a CYGNSS dataset
553
    def RMS ratio index(complete set):
      RMS ratio array = []
554
      sample array =[]
555
556
      ddm array=[]
      for item in range(0,complete set['sample'].size):
557
       for items in range(0,complete set['ddm'].size):
558
           RMS ratio array.append(root square ratio(complete set, item, items))
559
           sample array.append(complete set.isel(sample = item)['sample'].values)
560
           ddm array.append(complete set.isel(sample =item,ddm = items)['ddm'].values)
561
562
563
       sample array = np.array(sample array)
564
      ddm array = np.array(ddm array)
      RMS ratio array = np.array(RMS ratio array)
565
566
567
      RMS index frame = pd.DataFrame(sample array, columns = ['sample'])
      se = pd.Series(ddm array)
568
569
      RMS index frame['ddm'] = se.values
      se2 = pd.Series(RMS ratio array)
570
      RMS index frame['Highest/RMS'] = se2.values
571
      return RMS index frame
572
573
574 ##define a function that integrates the RMS ratio values for each DDM as a column in t
    def return ratio set(complete set, ratio index1):
575
      ratio index1a = ratio index1.drop(columns = ['Highest/RMS'])
576
577
      mindx = pd.MultiIndex.from frame(ratio index1a)
      ratio index1b = ratio index1.drop(columns = ['sample', 'ddm'])
578
579
      ratio array1 = np.array(ratio index1b)
      ratio index2 = pd.DataFrame(ratio array1, columns = ['RMS ratio'], index = mindx)
580
581
      ratio set = ratio index2.to xarray()
582
      complete set = xr.combine by coords([complete set, ratio set])
583
      return complete set
584
    ##define a function that will return an array of all maximum template matching coeffic
585
    def match coeff array(complete set, template image):
586
      coeff array =[]
587
      for item in range(0,complete set['sample'].size):
588
589
         for items in range(0,complete set['ddm'].size):
590
           samp select = item
591
          ddm select = items
592
           prepare test image(complete set,samp select, ddm select)
593
           testing image = cv2.imread('image to test.jpg',0)
594
595
          w, h = template image.shape[::-1]
596
597
           img = testing image.copy()
           method = eval('cv2.TM CCOEFF NORMED')
598
599
           # Apply template Matching
600
           res = cv2.matchTemplate(img,template image,method)
           min val, max val, min loc, max loc = cv2.minMaxLoc(res)
601
602
           coeff array.append(max val)
603
           plt.close()
```

```
604
605
           !rm image_to_test.jpg
      return coeff array
606
607
608 ##Define a function that prepares the test image for each template matching
    def prepare test image(complete set, sample selection, ddm selection):
609
610
      complete set.isel(sample = sample selection, ddm = ddm selection)['raw counts'].plot
611
      ax = plt.qca()
      ax.axes.xaxis.set visible(False)
612
613
      ax.axes.yaxis.set visible(False)
      plt.savefig('image to test.jpg', bbox inches = 'tight')
614
615
616 ### define a function to create a dataframe with multiindex and then incorporates Max
617 ##the original xarray dataset
618 def create complete with maxes(complete set, max coeff frame):
619
      sample array =[]
620
      ddm array=[]
621
      for item in range(0,complete set['sample'].size):
622
        for items in range(0,complete set['ddm'].size):
623
          sample_array.append(complete_set.isel(sample = item)['sample'].values)
          ddm array.append(complete set.isel(sample = item,ddm = items)['ddm'].values)
624
625
626
      sample array = np.array(sample array)
627
      ddm array = np.array(ddm array)
628
629
      multiIndex frame = pd.DataFrame(sample array, columns = ['sample'])
630
      se = pd.Series(ddm array)
631
      multiIndex frame['ddm'] = se.values
      max coeff frame1 = np.array(max_coeff_frame)
632
633
      mindx = pd.MultiIndex.from frame(multiIndex frame)
      multi frame var = pd.DataFrame(max coeff frame1, columns = ['Max Matching Coeff'], i
634
635
      set_with_maxes = multi_frame_var.to_xarray()
636
      complete set = xr.combine by coords([complete set, set with maxes])
637
638
      return complete_set
639
640 #define a function that converts wind speed components to wind speed with pythagorean
641 def create speed var(complete set):
      u array = complete set['wU10m'].values
642
643
      u array = u array.flatten()
      v array = complete set['wV10m'].values
644
      v array = v array.flatten()
645
646
      u array = np.square(u array)
      v array = np.square(v array)
647
648
      sum array = np.add(u array, v array)
649
      speed array = np.sqrt(sum array)
650
651
      sample_array =[]
652
      ddm array=[]
      for item in range(0,complete_set['sample'].size):
653
654
        for items in range(0,complete set['ddm'].size):
           sample array.append(complete set.isel(sample = item)['sample'].values)
655
656
          ddm array.append(complete set.isel(sample = item, ddm = items)['ddm'].values)
657
```

```
658
       sample array = np.array(sample array)
659
      ddm_array = np.array(ddm_array)
660
661
      multiIndex frame = pd.DataFrame(sample array, columns = ['sample'])
662
      se = pd.Series(ddm array)
      multiIndex frame['ddm'] = se.values
663
664
      mindx = pd.MultiIndex.from frame(multiIndex frame)
      multi_frame_var = pd.DataFrame(speed_array, columns = ['wind speed'], index = mindx)
665
666
      set_with_speed = multi_frame_var.to_xarray()
667
      complete set = xr.combine by coords([complete set, set with speed])
668
      return complete set
669
      ##create a function that will perform template matching; code source: Open cv2 pytho
670
671 def create matching(template image, testing image):
672
      methods = ['cv2.TM_CCOEFF', 'cv2.TM_CCOEFF_NORMED', 'cv2.TM_CCORR',
                 'cv2.TM_CCORR_NORMED', 'cv2.TM_SQDIFF', 'cv2.TM_SQDIFF_NORMED']
673
674
      w, h = template_image.shape[::-1]
675
676
      for meth in methods:
677
        img = testing_image.copy()
678
        method = eval(meth)
679
        # Apply template Matching
680
        res = cv2.matchTemplate(img,template image,method)
681
        min_val, max_val, min_loc, max_loc = cv2.minMaxLoc(res)
682
        # If the method is TM SQDIFF or TM SQDIFF NORMED, take minimum
683
        if method in [cv2.TM SQDIFF, cv2.TM SQDIFF NORMED]:
684
            top left = min_loc
685
686
        else:
             top_left = max_loc
687
688
        bottom_right = (top_left[0] + w, top_left[1] + h)
689
690
        cv2.rectangle(img,top left, bottom right, 0, 5)
691
        plt.subplot(121),plt.imshow(res,cmap = 'gray')
692
693
        plt.title('Matching Result'), plt.xticks([]), plt.yticks([])
694
        plt.subplot(122),plt.imshow(img,cmap = 'gray')
        plt.title('Detected Point'), plt.xticks([]), plt.yticks([])
695
696
        plt.suptitle(meth)
697
698
        plt.show()
699
700
        print(max val)
701
702 ###FUNCTIONS FOR EXPLORATORY ANALYSIS OF MODELING DATASET PHASE:
703
704 ##define a function that compares all relevant DDM calibration variables and wind spee
705 def full_scatter_compare(modeling_set):
706
      dset = modeling set
707
      array_list = [dset['ddm_average'], dset['RMS ratio'], dset['Max Matching Coeff'], ds
708
      array list = [item.values for item in array list]
709
      array list = [i.flatten() for i in array list]
710
      array_list[6] = np.sign((array_list[6]))*np.log(np.absolute(array_list[6])+1)
711
      array_list[5] = np.sign(array_list[5])*np.log(np.absolute(array_list[5])+1)
```

```
712
      array list[0] = np.log(array list[0])
      array_list[1] = np.log(array_list[1])
713
      names = ['ddm average', 'RMS ratio', 'Matching Coeff', 'wind speed', 'wave height',
714
      frame = pd.DataFrame.from dict(dict(zip(names, array list)))
715
      times array = modeling set['ddm timestamp utc'].values
716
      times array = np.repeat(times array, 4)
717
718
      pd.plotting.scatter matrix(frame, c = times array, cmap = 'viridis', figsize = (10,1
719
      plt.tight layout()
      plt.show()
720
721
722 #define a function that plots a box plot for a variable of interest, along with that v
723 def ddm box plot(ultimate set, var of interest):
724
      ddm array = ultimate set[var of interest].values
      ddm array = ddm array.flatten()
725
726
      ddm_array = ddm_array[~np.isnan(ddm_array)]
      plt.subplot(1,2,1)
727
728
      plt.boxplot(ddm array)
729
      plt.axis(ymin = 0)
      plt.title(var of interest)
730
731
      plt.subplot(1,2,2)
      if (var of interest == 'nbrcs') or (var of interest == 'les') or (var of interest ==
732
        plt.hist(ddm array,log = True, bins = 25)
733
734
        plt.hist(ddm array,bins = 25)
735
736
      plt.title(var of interest)
737
738 #define a function that allows the notebook user to select sample and DDM channel from
    ###corresponding callibration for the DDM or collocated wind speed/wave height associa
739
740 def ddm plots with vars(full dataset, sample sel, ddm sel):
741
      ddm = full dataset.sel(sample = sample sel, ddm = ddm sel)
742
      ddm['raw counts'].plot()
      samp = ddm['sample'].values
743
744
      chan = ddm['ddm'].values
      plt.title(f'Sample:{samp}
                                         DDM Channel:{chan}')
745
746
      variable_list = [ddm['ddm_average'], ddm['RMS ratio'], ddm['Max Matching Coeff'], dd
      variable list = [item.values for item in variable list]
747
      variable list = [[j.tolist()] for j in variable list]
748
      names = ['ddm average: ', 'RMS ratio: ', 'Max Matching Coeff: ', 'wind speed(m/s): '
749
      dictionary = dict(zip(names, variable list))
750
      frame = pd.DataFrame.from dict(dictionary)
751
752
      pd.set option('display.max columns', None)
      pd.set option('expand frame repr', False)
753
754
      frame = frame.round(3)
755
      print(frame)
      print("""
756
      """)
757
758
759 ##define a function that will zoom in on the points in a particular scatter plot, allo
760 ###also, display the correlation coefficient between the plotted variables, for the us
    def close_up_scatter(total_set, x_min, x_max, y_min, y_max, var_of_intA, var_of_intB):
761
762
      var list = [total set[var of intA].values, total set[var of intB].values]
      var list = [item.flatten() for item in var list]
763
      cmap = matplotlib.cm.get cmap("viridis", 5)
764
765
      times array = total set['ddm timestamp utc'].values
```

```
times array = np.repeat(times_array, 4)
766
767
      plt.figure(figsize = [8,8])
      ax1 = plt.scatter(x = var list[0], y = var list[1], c = times array, cmap = cmap, s
768
769
      plt.axis(xmin = x min, xmax = x max, ymin = y min, ymax = y max)
770
      cb = plt.colorbar(ax1)
      tick locator = ticker.MaxNLocator(nbins= 5)
771
772
      cb.locator = tick locator
773
      cb.update ticks()
774
      cb.ax.set yticklabels(['dropped string', 'March 15-April 15', 'April 16-May15', 'May
775
      plt.xlabel(var of intA)
776
      plt.ylabel(var of intB)
      corr = np.corrcoef(var list[0], var list[1])[0,1]
777
      print(f'Pearson Correlation Coefficient: {corr}')
778
779
780 ##define a function that will produce a correlation matrix for all calibration/wind/wa
781 def correlation matrix(modeling set):
782
      dset = modeling set
      array list = [dset['ddm average'], dset['RMS ratio'], dset['Max Matching Coeff'], ds
783
      array list = [item.values for item in array list]
784
785
      array_list = [i.flatten() for i in array_list]
      names = ['ddm average', 'RMS ratio', 'Matching Coeff', 'wind speed', 'wave height',
786
      frame = pd.DataFrame.from dict(dict(zip(names, array list)))
787
788
      coeff matrix = frame.corr()
      coeff matrix = coeff matrix.round(3)
789
790
      return coeff matrix
791
792 #define a function that plots two variables on a twin y-axis plot, with time as the sh
793 def create twin plot(total set, start samp, end samp, channel, var of intA, var of int
      partial set = total set.where((start samp <= total set['sample']) & (total set['samp</pre>
794
795
      fig, ax1 = plt.subplots()
796
      ax1.plot(partial set['ddm timestamp utc'],partial set[var of intA], '.', color = 'ta
797
      ax1.set xlabel('timestamp')
798
      ax1.set ylabel(var of intA, color = 'tab:red')
799
      ax2 = ax1.twinx()
      ax2.plot(partial_set['ddm_timestamp_utc'], partial_set[var_of_intB], '.', color = 't
800
      ax2.set ylabel(var of intB, color = 'tab:blue')
801
      plt.setp(ax1.get xticklabels(), rotation=30, horizontalalignment='right')
802
803
      start time = partial set['ddm timestamp utc'].values[0]
      end_time = partial_set['ddm_timestamp_utc'].values[partial set['ddm timestamp utc'].
804
805
      plt.title(f'Sampling Interval:{start time} to {end time}')
806
807 #define a function that will calculate and display the basic statistical summary for a
808
    def stat summaries(ultimate set, var of int):
      var array = ultimate set[var of int].values
809
810
811
      my mean = var array.mean()
812
      my median = np.ma.median(var array)
      my_std = var_array.std()
813
814
      my max = var array.max()
815
      my min = var array.min()
816
      my array = np.array([my mean, my median, my std, my max, my min])
      name_array = np.array(['Mean', 'Median', 'Std', 'Max', 'Min'])
817
818
819
      stat frame = pd.DataFrame(name array, columns = ['Statistics'])
```

```
stat frame['Values'] = my array
820
821
      stat_frame['Values'] = stat_frame['Values'].astype('float64')
      stat frame['Values'] = stat frame['Values'].round(3)
822
      print(stat frame)
823
824
825 ###FUNCTIONS FOR LINEAR MODELING PHASE:
826
827 #define a function that will take the variables of interest in our dataset and work th
828
    def create dataframe(modeling set):
829
      dset = modeling set
      array list = [dset['ddm average'], dset['RMS ratio'], dset['Max Matching Coeff'], ds
830
      array list = [item.values for item in array list]
831
      array list = [i.flatten() for i in array list]
832
      names = ['ddm average', 'RMS ratio', 'Matching Coeff', 'wind speed', 'wave height',
833
834
     frame = pd.DataFrame.from_dict(dict(zip(names, array_list)))
835
      return frame
836
837 #define a function that will take in a dataframe and remove variables as we find probl
838 def remove colinear var(modeling frame, var to remove):
839
      modeling_frame = modeling_frame.drop([var_to_remove], axis = 1)
      return modeling frame
840
841
842 ##define a funtion that will plot a fitted vs. residuals plot for linear model
843 def FitvResid(regress, X, y):
844
      dataframe = pd.concat([X,y], axis = 1)
      model fitted y = regress.fittedvalues
845
      plot lm = plt.figure()
846
      plot lm.axes[0] = sns.residplot(model fitted y, dataframe.columns[-1], data = datafr
847
      plot lm.axes[0].set title('Residuals vs. Fitted')
848
      plot lm.axes[0].set xlabel('Fitted values')
849
      plot lm.axes[0].set ylabel('Residuals')
850
851
852 #define a function that will actually calculate the VIF values given a modeling datafr
853 ##NOTE: much of this code is adapted from the page 'Detecting Multicolinearity with VI
854 def find_VIF(modeling_frame, dependent_var):
855
      X = modeling frame.drop([dependent var], axis = 1)
      vif data = pd.DataFrame()
856
      vif data["feature"] = X.columns
857
      vif data["VIF"] = [variance inflation factor(X.values, i) for i in range(len(X.colum
858
      vif data["VIF"] = vif data["VIF"].round(decimals = 3)
859
      return vif data
860
861
862 #define a function that will take in a dataframe of variables to be used in linear mod
863 ##and then calculate the RSS (residual sum of squares) for the model
864 ##NOTE: Much of this code was adapted from science.smith.edu
865 def mod subset(variable set):
866
     mod = sm.OLS(y,X[list(variable set)])
     regress = mod.fit()
867
868
     RSS = ((regress.predict(X[list(variable set)]) - y) ** 2).sum()
      dic = {'model':regress, 'RSS': RSS}
869
     return dic
870
871
872 #define a function that calls mod subset for each combination of regressor variables a
873 #NOTE: Much of this code was adapted from science.smith.edu
```

```
874 def highest RSS(num of regressors):
875
      array = []
      for combination in itertools.combinations(X.columns, num of regressors):
876
877
        array.append(mod subset(combination))
      models frame = pd.DataFrame(array)
878
      best mod = models frame.loc(models frame('RSS').argmin())
879
880
      return best mod
881
882 #define a function that plots cooks distances from our linear models
883 def cooks distances plot(regression mod):
      inf = influence(regress)
884
      C, P = inf.cooks distance
885
      , ax = plt.subplots(figsize=(9,6))
886
      ax.stem(C, markerfmt=",")
887
888
     ax.set xlabel("instance")
     ax.set ylabel("distance")
889
     ax.set title(f"Cook's Dist. Influentials Plot: {y.name} Model")
890
891
892 #define a function that plots fitted values against observed values for linear model
893 def fitVsobserved(results):
      fig, ax1 = plt.subplots(2,2)
894
895
      sm.graphics.plot fit(results, 0, ax = ax1[0,0])
896
      sm.graphics.plot fit(results, 1, ax = ax1[0,1])
      sm.graphics.plot fit(results, 2, ax = ax1[1,0])
897
898
      sm.graphics.plot fit(results, 3, ax = ax1[1,1])
      plt.tight layout()
899
     fig.set size inches(8,8)
900
901
902 #define a function that plots component plus residuals plot grid for variables in a mo
903 def ccpr plots(results):
     fig, ax1 = plt.subplots(2,2)
904
      sm.graphics.plot_ccpr(results, 0, ax = ax1[0,0])
905
906
      sm.graphics.plot ccpr(results, 1, ax = ax1[0,1])
      sm.graphics.plot ccpr(results, 2, ax = ax1[1,0])
907
      sm.graphics.plot_ccpr(results, 3, ax = ax1[1,1])
908
     plt.tight layout()
909
      fig.set size inches(8,8)
910
911
912 #define a function that presents outliers from bonferroni test
913 def bonf outlier(bonf test):
      bonf outliers = bonf test.where(bonf test['student resid'] > bonf test['bonf(p)'])
914
      bonf outliers = bonf outliers.dropna()
915
916
      bonf outliers = bonf outliers.astype('float64')
      bonf outliers = bonf outliers.round(3)
917
      return bonf outliers
918
919
920 #define function that plots a histogram of model residuals
921 def resid Hist(regress):
     mod resid = regress.resid
922
     fig, ax = plt.subplots(figsize =(10, 7))
923
     ax.hist(mod resid)
924
     ax.set xlabel('error')
925
      ax.set title('Residuals Distribution')
926
927
      plt.show()
```

```
928
      del mod resid
929
930 #define a function that compares averages for dependent variables of total modeling se
931
    def compare dependent average(outlier select, modeling set):
      av wind outliers = outlier select['wind speed'].values.mean()
932
      av wave outliers = outlier select['wSWH'].values.mean()
933
934
      av wind total = modeling set['wind speed'].values.mean()
      av wave total = modeling set['wSWH'].values.mean()
935
      print(f''' Average Wind Speed (Sample Subset): {av wind outliers.round(3)}
936
      Average Wave Height (Sample Subset): {av wave outliers.round(3)}
937
      Average Wind Speed (Total Set): {av wind total.round(3)}
938
      Average Wave Height (Total Set): {av wave total.round(3)}''')
939
940
941 ###FUNCTIONS FOR MACHINE LEARNING MODELING PHASE:
942
943 def ML data prep(ds, sample=False, size = 1000, replace=False, ext = ''):
944
945
      Input:
      ds = xarray dataset
946
947
      sample = If True, sample a subset from the full data set
      size = this will determine the size of the sample (only if sample = True)
948
      replace = If True, sample with replacement (only if sample = True)
949
950
      ext = an optional string added to allow for generating and saving multiple
             sampling data sets easily
951
952
953
      Output data frame ready for machine learning
      1 1 1
954
955
      # Removing uneeded variables and converting to dataframe
956
957
      dset = ds
      array list = [dset['ddm average'], dset['RMS ratio'], dset['Max Matching Coeff'], ds
958
959
      array list = [item.values for item in array list]
960
      array list = [i.flatten() for i in array list]
      names = ['ddm average', 'RMS ratio', 'Matching Coeff', 'wind speed', 'wSWH', 'nbrcs'
961
962
      df = pd.DataFrame.from_dict(dict(zip(names, array_list)))
963
964
      # Categorizing wind speed
965
      wind category = []
      n \text{ wind} = len(df)
966
967
      mean wind = df['wind speed'].mean()
      sd wind = df['wind speed'].std()
968
969
970
      for i in tqdm(range(0, n wind)):
         if df['wind speed'][i] < mean wind - sd wind:
971
           temp = 'Calm'
972
973
        elif df['wind speed'][i] > mean wind + sd wind:
          temp = 'Strong'
974
975
        else:
          temp = 'Mild'
976
977
978
        wind category.append(temp)
979
980
      df['wind category'] = wind category
981
```

```
982
      # Categorizing significant wave height
983
      wave category = []
      n wave = len(df)
984
985
      mean wave = df['wSWH'].mean()
      sd wave = df['wSWH'].std()
986
987
988
      for i in tqdm(range(0, n wave)):
989
        if df['wSWH'][i] < mean wave - sd wave:</pre>
          temp = 'Low'
990
991
        elif df['wSWH'][i] > mean wave + sd wave:
          temp = 'High'
992
993
        else:
         temp = 'Medium'
994
995
996
        wave_category.append(temp)
997
998
      df['wave_category'] = wave_category
999
1000
      if sample == True:
1001
        df.to pickle(f'{pathTeam}ML_data_sample{ext}.pkl')
1002
      else:
1003
         df.to pickle(f'{pathTeam}ML data{ext}.pkl')
1004
1005 def nclass classification mosaic plot(n classes, results):
1006
1007
        build a mosaic plot from the results of a classification
1008
1009
       parameters:
1010
       n classes: number of classes
1011
       results: results of the prediction in form of an array of arrays
1012
        In case of 3 classes the prdiction could look like
1013
1014
        [[10, 2, 4],
        [1, 12, 3],
1015
         [2, 2, 9]
1016
1017
        where there is one array for each class and each array holds the
1018
1019
        predictions for each class [class 1, class 2, class 3].
1020
        This is just a prototype including colors for 6 classes.
1021
        11 11 11
1022
        class lists = [range(n classes)]*2
1023
1024
        mosaic tuples = tuple(itertools.product(*class lists))
1025
1026
       res list = results[0]
1027
        for i, l in enumerate(results):
             if i == 0:
1028
1029
                 pass
1030
             else:
                 tmp = deque(1)
1031
                 tmp.rotate(-i)
1032
1033
                 res list.extend(tmp)
1034
        data = {t:res_list[i] for i,t in enumerate(mosaic_tuples)}
1035
```

```
1036
        fig, ax = plt.subplots(figsize=(8, 7))
1037
         plt.rcParams.update({'font.size': 16})
1038
1039
        font color = '#2c3e50'
1040
        # pallet = [
1041
               '#6a89cc',
1042
               '#4a69bd',
               '#1e3799',
1043
               '#0c2461',
1044
        #
1045
               '#82ccdd',
               '#60a3bc',
1046
        #
1047
        pallet = ["#1abc9c", "#3498db", "#e74c3c", "#f39c12", "#95a5a6"]
1048
        #pallet = ["#487eb0","#6a89cc", "#81cfe0","#00b5cc","#52b3d9"]
1049
1050
        colors = deque(pallet[:n_classes])
1051
        all colors = []
1052
        for i in range(n classes):
1053
             if i > 0:
1054
                 colors.rotate(-1)
1055
             all_colors.extend(colors)
1056
1057
        props = {(str(a), str(b)):{'color':all colors[i]} for i,(a, b) in enumerate(mosaic
1058
        labelizer = lambda k: ''
1059
1060
        p = mosaic(data, labelizer=labelizer, properties=props, ax=ax)
1061
1062
1063
        title font dict = {
1064
             'fontsize': 15,
             'color' : font_color,
1065
1066
         }
1067
        axis_label_font_dict = {
1068
             'fontsize': 10,
             'color' : font color,
1069
1070
1071
         ax.tick_params(axis = "x", which = "both", bottom = False, top = False)
1072
        ax.axes.yaxis.set ticks([])
1073
        ax.tick params(axis='x', which='major', labelsize=14)
1074
1075
        ax.set title('Mosaic Plot of Confusion Matrix', fontdict=title font dict, pad=25)
1076
         ax.set xlabel('Observed Class', fontdict=axis label font dict, labelpad=10)
1077
1078
         ax.set ylabel('Predicted Class', fontdict=axis label font dict, labelpad=35)
1079
         legend elements = [Patch(facecolor=all colors[i], label='Class {}'.format(i)) for
1080
         ax.legend(handles=legend elements, bbox to anchor=(1,1.018), fontsize=16)
1081
1082
1083
        plt.tight_layout()
1084
        plt.show()
```

Results and Discussion

Exploratory Data Analysis

CYGNSS data

Exploring an Initial CYGNSS Data Set (Associated with April 11)**

```
1 #read in the first CYGNSS file from google drive
2 ##for notebook user without access to drive, load in 'cyg firstfile.nc' from the 'Files
3 pathTeam = cwd + '/drive/My Drive/'
4 if os.path.exists(pathTeam + pathProfessor):
5 pathTeam += pathProfessor
6 pathTeam += Ben path # Should be a shortcut (Links to an external site.) to Team's shar
7 os.listdir(pathTeam)
   ['cyg firstfile.nc',
     'all_data_CYGNSS_0411.nc4',
     'all data CYGNSS 0411B.nc4',
     'ddm screenshot.png',
     'CYGNSS Background Collocated 20210411.nc',
     'set_4_11_RMS_av.nc',
     'modeling dataset.nc'
     'wValues 20210411.pkl']
1 cyg_data_set = xr.open_dataset(f'{pathTeam}cyg_firstfile.nc')
2 cyg data set
   xarray.Dataset
                       (ddm: 4, delay: 128, doppler: 20, sample: 2416)
   ▶ Dimensions:
   ▼ Coordinates:
      sample
                       (sample)
      ddm
                       (ddm)
      ddm_timestamp... (sample)
      sp lat
                       (sample, ddm)
      sp lon
                       (sample, ddm)
    ▼ Data variables:
      spacecraft_id
                       (sample)
      spacecraft_num
                       (sample)
      ddm_sample_in...
                       (sample)
                       (sample, ddm)
      prn_code
      raw_counts
                       (sample, ddm, delay, doppler)
                                                       float64 ...
    ► Attributes: (28)
```

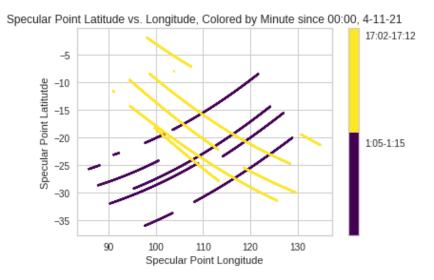
As the summary for the dataset shows, this is a four-dimensional delay-doppler map record across 2416 samples taken on one day: April 11, 2021. Next, we dig into the dataset to understand its structure and

what it shows about the CYGNSS GNSS scatterometry process.

Dissecting the Structure of a CYGNSS Dataset:

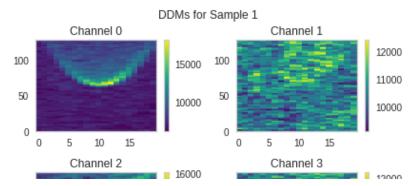
First, we looked at some plots of specular point latitude against specular point longitude to get an idea of how the satellites are moving over the sampling time intervals for this dataset (samples taken on April 11, 2021).

```
1 #call the function that will plot specular point latitude against specular point longit
2 #Note: this function allows the user to specify which sampling interval to plot, by sta
3 ###in this cell, we will just plot all the samples across all timestamps and DDMs
4 start_sample_index = 0
5 end_sample_index = 2415
6 location_time_plot(cyg_data_set, start_sample_index, end_sample_index)
```

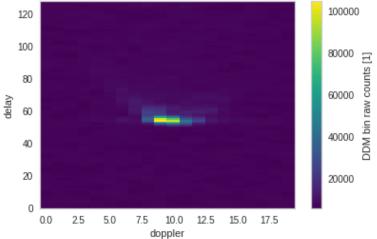


Here we can see that two CYGNSS satellites took Delay Doppler Map measurements across two time intervals, each logging four Delay Doppler Maps (from four specular points) every half second for each of those two intervals. Now, we examine four of the thousands of Delay Doppler Maps in the CYGNSS single file dataset, to get a sense of how different they can be to one another.

```
1 ##looking at the DDMs for sample index 1:
2 sample_in = 1
3 ddm_plots(cyg_data_set, sample_in)
4 del sample_in
```



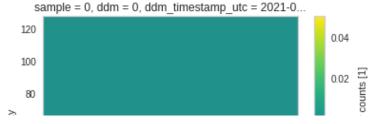
It's clear that the DDM can give a few broad types of images. One is the well-defined parabola, as seen in the DDM for channel 0 above. Another is the less well-defined parabola. The third is a more scrambled image, as seen in the DDMs for channels 1 and 3 in this sample. The last, not seen in this particular sample, is the occurrence of a single, bright specular point, surrounded by dark blue, as with this DDM:



Interestingly, there are a few unusuable Delay Doppler Maps recorded by the satellites, usually at the beginning of major sampling time intervals. In this set, we have four such DDMs at the very start of the dataset (sample 0), and another four at sample 1206, the start of the second sampling/time interval for the day. We look at the one from sample 0:

```
1 cyg_data_set.sel(sample=0, ddm=0)['raw_counts'].plot()
```

<matplotlib.collections.QuadMesh at 0x7f86d21e0490>



Further down, this dataset is processed through several functions that remove different types of junk/missing data, this type included.

ECWMF data

aoppier

Exploring the ECWMF Wind Speed/Wave Height Background Files (associated with Aprill 11)

```
1 # Importing netCDF files with background wind speed data for April 11, 2021
2 ##first, must reset pathTeam to exclude Ben_path information:
3 pathTeam = cwd + '/drive/My Drive/'
4 ##Check to add professor's path
5 if os.path.exists(pathTeam + pathProfessor):
6 pathTeam += pathProfessor
7 pathTeam += David path # Should be a shortcut (Links to an external site.) to Team's sh
8 os.listdir(pathTeam)
   ['cyg firstfile sps.pkl',
    'cyg.ddmi.s20210411-010506-e20210411-171248.ll.power-brcs-full.a30.d31.nc',
    'ecmwf.t00z.pgrb.0p125.f000 2021041100.nc',
    'ecmwf.t12z.pgrb.0p125.f000 2021041112.nc',
    'ecmwf.t18z.pgrb.0p125.f000 2021031118.nc',
    'CYGNSS 0311.pkl',
    'CYGNSS 0411.pkl',
    'CYGNSS Background Collocated 20210311.nc',
    'modeling dataset.nc',
    'wValues 20210311.pkl',
    'ML_data_sample2.pkl',
    'wValues 20210411.pkl',
    'wValues 20210411 sample.pkl',
    'ML data.pkl']
1 # Importing netCDF files with background wind speed data for April 11, 2021
2 # Adding speed column to dataset
3 ds00 = xr.open dataset(f'{pathTeam}ecmwf.t00z.pgrb.0p125.f000 2021041100.nc')
4 ds00 = ds00.assign(SPD = np.sqrt(ds00.U10m**2 + ds00.V10m**2)) # Calculates wind speed
5 ds00.info
   <bound method Dataset.info of <xarray.Dataset>
   Dimensions: (x: 1441, y: 2880)
   Coordinates:
               (x) float32 ...
       lat.
               (y) float32 ...
   Dimensions without coordinates: x, y
```

This dataset is a 2D dataset with the zonal and meridonal wind speed vectors and significant wave height values for every latitude and longitude pair in 1/8 incriments on April 11, 2021. The wind speed variable was added in the notebook because wind speed can be useful for visual data analysis.

The team coded a function that generates 5 or 6 plots (depending on the boolean passed in the function parameter 'overlay') for a region surrounding input coordinates.

Plot 1: Shows a U10m vs V10m scatter plot and prints the correlation between the two variables in the plot title

Plot 2: Shows a plot with the ECWMF grid layout in blue and the input coordinates as a red dot. The distance from the input corrdinates to the nearest grid point is printed in the title

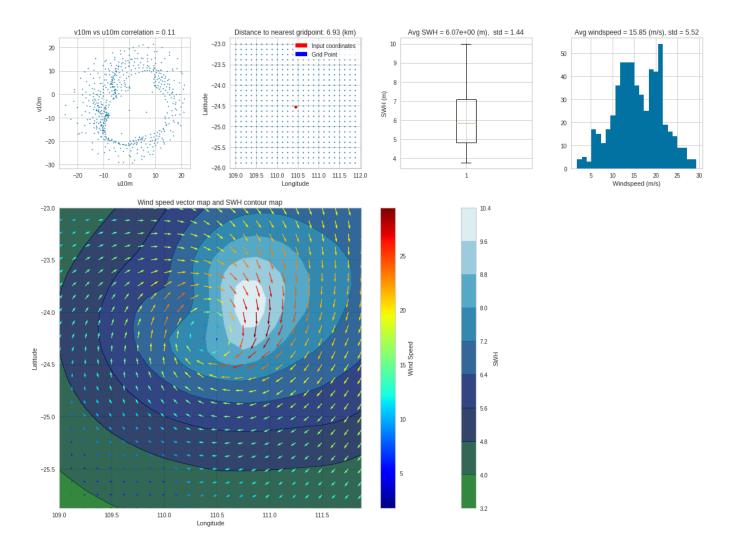
Plot 3: Shows a significant wave height boxplot. The title has the mean and standard deviation of the SWH values in the region.

Plot 4: Shows a wind speed histogram. The title has the mean and standard deviation of the wind speed in the region.

Plot 5: Shows a vector map of wind speed plotted over a contour map for SWH for the region. By changing the value passed in the function through the parameter 'alpha = ', you can adjust the transparancy of the SWH contour plot. The value must be in the range (0, 1]. Based on the pattern shows, it appears like this area is over a spiraling wind pattern and that wind patter has caused the SWH to be higher near the center of the wind pattern.

Note: If 'overlay = True' is passed into the function, Plot 5 splits into two plots and displays the wind speed vector map and SWH contour map seperately.

```
1 vizualize region(ds00, -24.53, 110.44, alpha=0.8)
```



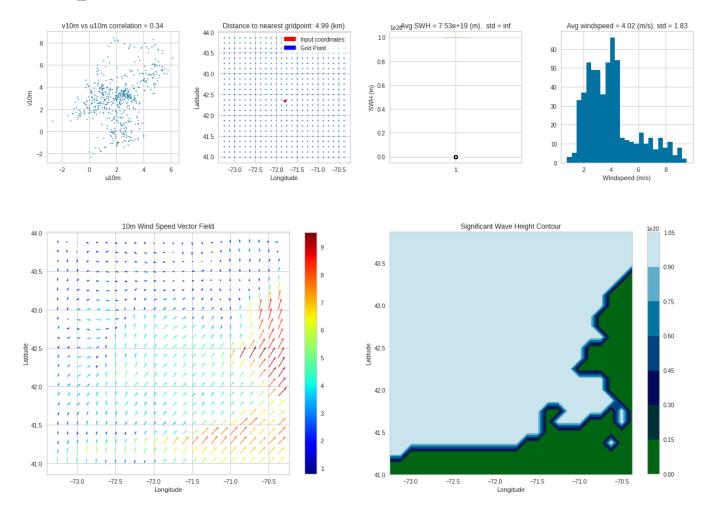
Plot 1: The correlation between the V10m and U10m variables for this region is 0.11. This indicates that there is not a strong correlation between the two variables in this region.

Plot 2: Shows where the input coordinates are in relation to the background grid. The distance from the input coordinates to the nearest grid point is 6.93 km.

Plot 3: The average SWH for this region is 6.07 meters with a standard deviation of 1.44. The median of this boxplot appears to be a little under 6. The mean is higher than the median which indicates that the data is positively skewed.

Plot 4: There is an average windspeed of 15.85 m/s with a standard deviation of 5.52 for the region. The histogram shows a bimodal distribution. It is possible that the wind vortex in the region (visualized in plot 5) is causing this bimodal distribution since the windspeeds act differently towards the center of the vortex when compared to the rest of the region.

Plot 5: The SWH contour and wind speed vector map show how these two variables interact with each other. This plot shows that the areas with higher SWH values tend to have a stronger wind blowing above. This also demonstrates that the 'eye of the storm' is fairly calm compared to its surrounding region as the wind speed vectors dramatically lower at the center of the vortex than they are in the area immediately surrounding.



Plot 1: The correlation between V10m and U10m for this region. Due to a difference in temperature and pressure between the ocean and land, wind tends to blow from a body of water to a land mass during the day, and from a land mass to a body of water at night. <u>Link</u>. The higher correlation in this region could be due in part to the pecense of a coastline (visualized in plot 6).

Plot 2: Shows where the input coordinates are in relation to the background grid. The distance from the input coordinates to the nearest grid point is 4.99 km.

Plot 3: The boxplot of SWH demonstrates something strange happening with the data, as the mean value is 7.54e19 m. The findings and how the team handled this is discussed further below.

Plot 4: The average windspeed for the region is 4.02 m/s with a standard deviation of 1.83. This histogram is unimodal and positively skewed.

Plot 5: The wind speed vector field shows calm winds in general. The bottom right area of the plot shows some stronger winds. This is likely due to the fact that the coastline is in that area of the region.

Plot 6: The SWH contour plot again shows an issue with the data as the color scale has a maximum value of 1e20. Visually, some conclusions can still be drawn. The difference in SWH is due to the fact that both land and water appear in the region. This plot shows the coastline present in the region. The water is colored green, and the land is colored light blue.

Note: This region might be more interesting to analyze with overley set to True, however, the team wanted to demonstrate that functionality so the region was analyzed with overlay set to False.

As shown in plots 3 and 6 of the visualization above, there are SWH values in the dataset with unreasonable values. What the team found was that ECWMF uses the value 1e20 in place of a Null value. This convention from ECWMF required action from the team before the interpolation process could begin. In order to interpolate the SWH values, the definition of the search neighborhood needed to be adjusted. If a specular point has a grid point in the serach neighborhood that is over land, there will be a 1e20 value recorded for SWH. To account for this, the team omitted all grid points in the search area that have a SWH value of 1e20 when interpolating SWH values. In the case that the search neighborhood for the specular point has no usable grid points for SWH, the value is set to np.nan per the request of the project mentor. Additional functionality was added to the interpolation function late in the process that allowed the team to track the proportion of observations with a SWH search neighborhood affected by this fact. The team tracked 84,042 observations over 5 days and found that around 15% of the observations had less than 4 neighbors in the search neighborhood for interpolating SWH. A summary of this process is shown in the table below.

| Value | 4 Neighbors | 3 Neighbors | 2 Neighbors | 1 Neighbors | 0 Neighbors |
|-------|-------------|-------------|-------------|-------------|-------------|
| U10m | 1.00 | 0 | 0 | 0 | 0 |
| V10m | 1.00 | 0 | 0 | 0 | 0 |
| SWH | 0.84 | 0.10 | 0.08 | 0.07 | 0.13 |

Data Collocation

First, the team needed to interpolate the Wind Speed/Wave Height Background Data on some geometric/location principle, so we could get just one wind speed and wave height value per DDM. This entailed getting a list of the timestamps and latitude/longitude of all specular points in the CYGNSS file extracted:

```
2 latlon_frame = collect_latlons(cyg_data_set)
```

This dataframe was then transfered between team members as a pickle file, for interpolation of wind speed/wave height data.

Interpolation

Interpolation of the Background Wind Grid Data in Advance of Collocation with the Original CYGNSS Dataset:

The first step for interpolation was to open and understand the structure of the timestamp pickle file that was generated above.

```
1 ds_cord = pd.read_pickle(f'{pathTeam}CYGNSS_0411.pkl')
2 ds cord
```

| | timestamp | sp_lat | sp_lon |
|------|-------------------------------|---------------------|--------------------|
| 0 | 2021-04-11 01:05:06.499261678 | -31.94055938720703 | 90.30609893798828 |
| 1 | 2021-04-11 01:05:06.499261678 | -25.70931053161621 | 85.83236694335938 |
| 2 | 2021-04-11 01:05:06.499261678 | -28.654094696044922 | 87.773681640625 |
| 3 | 2021-04-11 01:05:06.499261678 | -23.13581657409668 | 91.06610870361328 |
| 4 | 2021-04-11 01:05:06.999261612 | -31.931716918945312 | 90.33650970458984 |
| | | | |
| 9659 | 2021-04-11 17:12:47.999261690 | -24.73063087463379 | 128.42019653320312 |
| 9660 | 2021-04-11 17:12:48.499261605 | -29.96910858154297 | 129.6576385498047 |
| 9661 | 2021-04-11 17:12:48.499261605 | -21.37551498413086 | 134.84649658203125 |
| 9662 | 2021-04-11 17:12:48.499261605 | -31.438047409057617 | 125.68804931640625 |
| 9663 | 2021-04-11 17:12:48.499261605 | -24.740629196166992 | 128.4476318359375 |
| | | | |

9664 rows × 3 columns

The pickle file loads as a pandas dataframe with the variable's 'timestamp', 'sp_lat', and 'sp_lon.' The goal will be to interpolate U10m, V10m, and SWH based on the sp_lat and sp_lon varibales.

The next thing the team did in their efforts to interpolate the wind speed and wave height data for the CYGNSS specular point locations was to code a funcion that finds the closest ECMWF data collection point to a CYGNSS specular point. This function will work with any input latitude and longitude cordinates but was not used in that manner.

The function above returns a tuple with the U10m, V10m, and SWH values of the ECWMF data collection point closest to the first specular point in the CYGNSS data set for April 11th, 2021. While this is not the most accurate way to assign these values to the specular point locations, this was useful in motivating the function used to interpolate the desired data for each specular point.

The team did not get all the ECMWF datasets needed initially. In order to save time, the team coded a function that will read the CYGNSS timestamp pickle file for a date and print out what ECMWF files are needed to run the interpolation function.

```
1 ecmwf_check(4, 11)
    Need ECMWF file: ecmwf.t00z.pgrb.0p125.f000_2021041100.nc
    Need ECMWF file: ecmwf.t12z.pgrb.0p125.f000_2021041112.nc
```

Once the team was able to confirm that the proper files were uploaded, they were prepared to interpolate the data for that day.

WARNING the function below takes around 3 minutes and 30 seconds to execute for April 11, 2021.

```
count 9664.000000 9664.000000 9664.000000 9664.000000 7824.000000
                                                     2.615569
std
       6.789308 10.319270
                               4.328611
                                           5.575101
                                                      0.699901
      -35.971329
                  85.832367
                              -7.593077 -13.718920
                                                      0.146045
min
25%
      -26.438673 101.445791
                              -2.585555
                                         -1.796945
                                                      2.293432
50%
      -21.540831 108.835327
                              0.975676
                                          1.727017
                                                      2.700940
75%
      -16.813571 118.226568
                               4.065569
                                          6.893683
                                                      2.964786
       -1.941648 134.846497 11.554374 12.180312
                                                      4.827818
max
```

```
Neighbor count for U10m:
4 9664
Name: U10m_neighbor, dtype: int64
Neighbor count for V10m:
4 9664
Name: V10m_neighbor, dtype: int64
Neighbor count for SWH:
0 1840
1 32
2 65
```

```
3 47
4 7680
Name: SWH neighbor, dtype: int64
```

Functionality was added to the code to allow for a subset of the data to be run through the interpolate_date function. By passing 'subset = True' into the function, only the first n observations will be interpolated. The value n is set to a default value of 500 but can be adjusted by passing 'len_subset = n' to the interpolate_date function. Interpolation of the first 500 observations of this dataset only takes around 11 seconds

```
1 interpolate_date(4, 11, subset = True)
  ds00 was loaded
  ds12 was loaded
  100% | 500/500 [00:12<00:00, 38.93it/s]
                         lon
               lat
                                  wU10m
                                             wV10m
                                                        wSWH
  count 500.000000 500.000000 500.000000 500.000000 500.000000
  mean -29.330124 92.031972 -3.290570 7.005655 2.693062
         3.723712 3.935408 3.345849
  std
                                         1.127885 0.222715
        -35.971329 85.832367 -7.041207 4.651071 2.232990
  min
  25%
       -31.555711 88.989885 -5.412166 6.222458 2.581645
       -28.650121 91.321507 -4.385120
  50%
                                         7.027007 2.750525
        -25.707332 94.006428 -2.959006
  75%
                                         7.537441
                                                     2.840100
                                         9.521032 3.014740
        -22.797634 100.089447
  max
                              5.582742
  Neighbor count for U10m:
      500
  Name: U10m neighbor, dtype: int64
  Neighbor count for V10m:
       500
  Name: V10m neighbor, dtype: int64
  Neighbor count for SWH:
       500
  Name: SWH neighbor, dtype: int64
```

The printout for this funcion has 3 sections. The first section shows the ECMWF files needed. In this case, the files loaded match the files needed as found by the 'ecmwf_check' function. The second part of the printout tracks the loop's progress. The third part of the printout a summary of the new dataset. This summary shows no strange or unexpected values, so the interpolation process proceeded as planned. This process was done for all 45 days of data provided to the team. As the datasets with the interpolated data were completed, the collocation process began.

```
1 del ds00, ds cord # Deleting uneeded global variales
```

Collocation

Joining the Interpolated Wind/Wave Data Back into the Original CYGNSS Dataset:

```
1 # Importing the interpolated wind data
2 ##reset pathTeam to exclude David path information
3 pathTeam = cwd + '/drive/My Drive/'
4 ##Check to add professor path
5 if os.path.exists(pathTeam + pathProfessor):
6 pathTeam += pathProfessor
7 pathTeam += Ben path # Should be a shortcut (Links to an external site.) to Team's shar
8 os.listdir(pathTeam)
   ['cyg_firstfile.nc',
    'all data CYGNSS 0411.nc4',
    'all data CYGNSS 0411B.nc4',
    'ddm screenshot.png',
    'CYGNSS Background Collocated 20210411.nc',
    'set 4 11 RMS av.nc',
    'modeling dataset.nc',
    'wValues 20210411.pkl']
1 #read in the dataframe of interpolated wind/wave values created in the previous section
2 coll set = pd.read pickle(f'{pathTeam}wValues 20210411.pkl')
```

All that remained was for the interpolated data to be joined back into the original xarray CYGNSS dataset:

```
1 #Now, we combined that interpolated data back into the original CYGNSS dataset
2 set_to_coll = cyg_data_set
3 collocated_set = integrate_sets(set_to_coll, coll_set)
4 collocated_set
```

xarray.Dataset

```
▶ Dimensions:
                      (ddm: 4, delay: 128, doppler: 20, sample: 2416)
▼ Coordinates:
   sample
                      (sample)
   ddm
                      (ddm)
   ddm_timestamp...
                      (sample)
   sp_lat
                      (sample, ddm)
   sp_lon
                      (sample, ddm)
▼ Data variables:
                                                           float32 249.0 249.0 249.0 ... 55.0 55.0
   spacecraft_id
                      (sample)
   spacecraft_num
                      (sample)
   ddm_sample_in...
                      (sample)
                                                           float64 7.812e+03 7.813e+03 ... 1.23...
   prn_code
                      (sample, ddm)
   raw_counts
                      (sample, ddm, delay, doppler)
   wU10m
                      (sample, ddm)
                                                           float64 -2.678 -5.944 ... -4.316 -4.25
                                                                                                    wV10m
                      (sample, ddm)
                                                           float64 6.087 7.058 6.24 ... -1.947 -1....
   wSWH
                      (sample, ddm)
                                                           float64 2.624 3.01 2.753 ... nan nan nan
► Attributes: (28)
```

So we see that we now have the original CYGNSS dataset, but this time with background wind/wave data, interpolated as a weighted average of wind/wave values around each specular point, all saved in their respective new variables.

Data Cleaning after Collocation

Now it remained for the team to remove 'junk' (all zero DDM) samples, as well as samples with significant wave height data that was simply missing (as associated specular point may be on land):

```
1 ##clean the collocated set of its 'junk' DDM samples:
2 collocated_clean = collocated_set.where(collocated_set['raw_counts']!= 0)
3 collocated_clean = collocated_clean.dropna(dim = 'sample')
4 collocated_clean
```

xarray.Dataset

```
▶ Dimensions:
                      (ddm: 4, delay: 128, doppler: 20, sample: 1354)
▼ Coordinates:
   sample
                      (sample)
   ddm
                      (ddm)
   ddm_timestamp... (sample)
   sp_lat
                      (sample, ddm)
   sp lon
                      (sample, ddm)
▼ Data variables:
   spacecraft_id
                      (sample, ddm, delay, doppler)
   spacecraft_num
                      (sample, ddm, delay, doppler)
                      (sample, ddm, delay, doppler)
                                                           float64 7.813e+03 7.813e+03 ... 1.23...
   ddm_sample_in...
   prn code
                      (sample, ddm, delay, doppler)
                      (sample, ddm, delay, doppler)
                                                           float64 6.974e+03 6.614e+03 ... 9.56...
   raw_counts
   wU10m
                      (sample, ddm, delay, doppler)
                                                           float64 -2.678 -2.678 ... 4.529 4.529
                                                                                                    wV10m
                      (sample, ddm, delay, doppler)
                                                           float64 6.087 6.087 ... -0.4586 -0.4586
                                                                                                    wSWH
                      (sample, ddm, delay, doppler)
                                                           float64 2.624 2.624 2.624 ... 2.092 2....
                                                                                                    ► Attributes: (28)
```

We can see that two samples have been removed for containing useless DDMs. Because the second code line in the previous cell removes all samples with any 'NaN' values, it eliminated all samples with junk DDMs or with any 'nan' values for significant wave height.

So it seems that for April 11, 2021, 1354 samples were retained with all clean and collocated data.

The team repeated the collocated and cleaning process for nearly every CYGNSS FULL DDM file NASA had available for samples taken from March 1 - Sep 1 of 2021. This collocated database is saved as a

collection of netCDF files, accessible through a google drive shortcut in the Spire project folder, under 'Spire_Clean_Collocated'.

The collocated/cleaned database contains 45 files in all, with the total amount of data retained after cleaning being 76.15% of the original, uncleaned CYGNSS data.

```
1 del collocated_clean
```

DDM Calibration

Once the team had managed to interpolate/collocate the background wind/wave data for each CYGNSS set of interest, we proceeded to process/find all the desired calibrations of the DDM data we would eventually seek to model with. This section demonstrates the processing of all those calibrations and their inclusion as variables in the greater dataset for just the date 4/11/21.

IMPORTANT NOTE: Even though we demonstrated the cleaning process on the collocated data we made available to Spire in the previous section, it is necessary for files we wanted to callibrate DDMs for (and model with) that all samples be retained in their original order from CYGNSS, if the NBRCS/LES retrieval function is to work properly. Hence, we start this section by reading in the original, uncleaned collocated 4/11/21 dataset, and then we clean it of junk DDMs and 'nan' significant wave height values later on in this section (after retrieval of NBRCS/LES values has been performed).

DDM Average Calculation

```
1 #first, we open the datafile that has collocated wind/wave data for 4/11
2 data_set_411 = collocated_set
3 del collocated set
```

First, we calculated the simple DDM average calibration- a simple average of raw counts values in a 10 x 5 area around the specular point bin of each DDM:

```
1 ##NOTE: This cell takes approximately 30 seconds to execute
2 sample_first = data_set_411.isel(sample=0)['sample']
3 ##the following values are to set the limits on delay and doppler for our 10x5 area ave
4 delay_start = 60
5 delay_end = 70
6 doppler_start = 8
7 doppler_end = 13
```

```
8 data_set_411 = DDM_averages(data_set_411,sample_first,delay_start,delay_end,doppler_sta
1 data_set_411
```

xarray.Dataset

```
▶ Dimensions:
                      (ddm: 4, delay: 128, doppler: 20, sample: 2416)
▼ Coordinates:
   sample
                      (sample)
   ddm
                      (ddm)
   ddm_timestamp... (sample)
   sp_lat
                      (sample, ddm)
   sp_lon
                      (sample, ddm)
▼ Data variables:
   ddm_average
                      (sample, ddm)
                                                            float64 0.0 0.0 0.0 ... 1.409e+04 6.56...
                                                            float32 249.0 249.0 249.0 ... 55.0 55.0
   spacecraft_id
                      (sample)
   spacecraft_num
                      (sample)
                                                            float64 7.812e+03 7.813e+03 ... 1.23...
   ddm_sample_in...
                      (sample)
   prn_code
                      (sample, ddm)
                                                            float64 0.0 0.0 0.0 ... 6.114e+03 6.03...
   raw_counts
                      (sample, ddm, delay, doppler)
                                                            float64 -2.678 -5.944 ... -4.316 -4.25
   wU10m
                      (sample, ddm)
                                                                                                    wV10m
                                                            float64 6.087 7.058 6.24 ... -1.947 -1....
                      (sample, ddm)
   wSWH
                      (sample, ddm)
                                                            float64 2.624 3.01 2.753 ... nan nan nan
► Attributes: (28)
```

We can see that the ddm averages have been calculated and added back into our dataset.

RMS Ratio Calculation

Next, we calculated the RMS ratio values (the highest power value for a given DDM divided by the Root Mean Square of the rest of the power values). Again, the team and instructor considered that it might be a useful statistic for eventual modeling:

```
1 ##calculate RMS ratio values for each DDM to include as a column in set_4_11
2 ###first, build RMS ratio index as a dataframe
3 ###then, combine content of dataframe back into CYGNSS dataset
4 ###NOTE: The warnings that arise as this cell executes occur when there are junk DDMs (
5 ###However, samples with these DDMs are removed further down in this calibration/data p
6 ###WARNING: takes a few minutes to run on a dataset with 2416 samples
7 ###NOTE: To test this function on a smaller set of data the user could first slice off
8 ###CONTINUED: And then process the smaller set through the functions
9 ###CONTINUED: Using the following 3 lines of code (here commented out):
10 ### data_set_partition = data_set_411.sel(sample = slice(starting_sample, ending_sample)
11 ### RMS Ratio index = RMS ratio index(data set partition)
```

12 ### data set partition = return ratio set(data set partition, RMS Ratio index)

13 RMS Ratio index = RMS ratio index(data set 411)

```
14 data set 411 = return ratio set(data set 411, RMS Ratio index)
15 data_set_411
     /usr/local/lib/python3.7/dist-packages/ipykernel launcher.py:544: RuntimeWarning: inv
     /usr/local/lib/python3.7/dist-packages/ipykernel launcher.py:544: RuntimeWarning: inv
     /usr/local/lib/python3.7/dist-packages/ipykernel launcher.py:544: RuntimeWarning: inv
     /usr/local/lib/python3.7/dist-packages/ipykernel_launcher.py:544: RuntimeWarning: inv
     /usr/local/lib/python3.7/dist-packages/ipykernel launcher.py:544: RuntimeWarning: inv
     /usr/local/lib/python3.7/dist-packages/ipykernel launcher.py:544: RuntimeWarning: inv
     /usr/local/lib/python3.7/dist-packages/ipykernel launcher.py:544: RuntimeWarning: inv
     xarray.Dataset
     ▶ Dimensions:
                         (ddm: 4, delay: 128, doppler: 20, sample: 2416)
     ▼ Coordinates:
        sample
                         (sample)
                                                            int64 0 1 2 3 4 ... 2412 2413 2414 2...
        ddm
                         (ddm)
        ddm_timestamp...
                         (sample)
        sp_lat
                         (sample, ddm)
        sp_lon
                         (sample, ddm)
     ▼ Data variables:
        RMS ratio
                         (sample, ddm)
                                                          float64 nan nan nan ... 1.209 3.871 1.03
        ddm_average
                         (sample, ddm)
                                                          float64 0.0 0.0 0.0 ... 1.409e+04 6.56...
        spacecraft_id
                         (sample)
        spacecraft num
                         (sample)
                         (sample)
        ddm sample in...
                         (sample, ddm)
        prn code
                                                          float64 0.0 0.0 0.0 ... 6.114e+03 6.03...
        raw counts
                         (sample, ddm, delay, doppler)
        wU10m
                                                          float64 -2.678 -5.944 ... -4.316 -4.25
                         (sample, ddm)
        wV10m
                                                          float64 6.087 7.058 6.24 ... -1.947 -1....
                         (sample, ddm)
        wSWH
                         (sample, ddm)
                                                          float64 2.624 3.01 2.753 ... nan nan nan
     ► Attributes: (28)
```

Note: The RMS ratio value was set equal to 'nan' by the function for the first (and 1206th) samples because the DDMs at this point of this section and at those samples are junk, with all power values still set to zero, so RMS ratio = 0 and division by zero is impossible. This doesn't really matter, as those samples will be dropped by a cleaning command further down. The only reason we didn't drop them already is, again, because it is necessary to retain all original samples until for the NBRCS/LES retrieval function to work properly.

NBRCS/LES Retrieval

Next, we found the NASA files containing NBRCS and LES values corresponding to these samples/this date, and worked NBRCS and LES into the dataset as well. NOTE: samples in our set were taken by two separate satellites (hence the two separate sampling intervals), so it became necessary to perform this

data match-up in two halves, to avoid writing a function that would need to search through 170,000 samples in the NASA file.

```
1 ###bring in NBRCS and LES values for these DDMs (from NASA's CYGNSS ALL DATA dataset)
2 ##start by reading in data from that larger CYGNSS Lv1 file
3 ##open file containing nbrcs data and les data
4 ##again this file can be found in the 'Files needed to run' folder
5 all data set = xr.open dataset(f'{pathTeam}all data CYGNSS 0411.nc4')
```

The first sampling interval runs from sample 0 to sample 1205, so we isolate those samples. In the NASA dataset, the corresponding samples run from sample indexes 7812 to 9017, so we isolate those samples and bring their NBRCS and LES values into the original set.

```
1 #extract first sampling interval's nbrcs/les values
2 first half set = NBRCS LES vals(all data set, data set 411, 7812, 9017, 0, 1205)
```

The second sampling interval runs from sample 1206 to sample 2415. In the NASA dataset, the corresponding samples run from sample indexes 122727 to 123936. We isolate those samples and bring their NBRCS and LES values into the original set. We must read in a new NASA file however, as there is a separate file for the different satellite.

```
1 #open the dataset for the second satellite for 4/11 since our dataset's second time int
2 all_data_set = xr.open_dataset(f'{pathTeam}all_data_CYGNSS_0411B.nc4')

1 second_half_set = NBRCS_LES_vals(all_data_set, data_set_411, 122727, 123936, 1206, 2415)

1 #combine the two half sets into one set, with original sample values and nbrcs/les incl
2 data_set_411 = xr.combine_by_coords([second_half_set, first_half_set])
3 data set 411
```

```
      ▶ Dimensions:
      (ddm: 4, delay: 128, doppler: 20, sample: 2416)

      ▼ Coordinates:
      sample
      (sample)
      int64
      0 1 2 3 4 ... 2412 2413 2414 2...

      ddm
      (ddm)
      int64
      0 1 2 3

      ddm_timestamp...
      (sample)
      datetime64[ns]
      2021-04-11T01:05:06.499261...

      sp_lat
      (sample, ddm)
      float32
      -31.94 -25.71 ... -31.44 -24.74

      sp_lon
      (sample, ddm)
      float32
      90.31 85.83 87.77 ... 125.7 12...
```

▼ Data variables:

Importantly, many of the NBRCS and LES values from the NASA set are 'nan'. NOTE: we did double check to make sure those values were given by NASA as 'nan' in the original set, so the 'nan' values there are not the result of any problem in our data match-up. However, it is useful to eliminate samples with any 'nan' values for NBRCS or LES:

Removing Samples with NaNs

```
wSWH (sample ddm) float64 2 624 3 01 2 753 nan nan nan D 1 ##remove all samples with 'nan' for nbrcs/les/wSWH
2 ##NOTE: This command also removed samples with junk DDMs (samples 0 and 1206), presumab 3 ##NASA did not calculate nbrcs/les for their all-zero DDMs.
4 data_set_411 = data_set_411.dropna(dim = 'sample')
5 data set 411
```

- 1 ##delete superfluous variables from the NBRCS/LES gathering process:
- 2 del first_half_set, second_half_set, all_data_set

sample

(sample)

int64 2 3 4 5 7 ... 1935 1980 1988 2...

Maximum Template Matching Coefficient Calculation

sp_iat (sample, dum)

110al32 -31.92 -23.09 ... -20.20 -20.13

The dataset now has all DDM calibrations but one: Maximum Template Matching Coefficient. Also, there are no missing values for 'nbrcs' or 'les'. Our work to perform the matching began with choosing an image for an 'ideal' ddm template image and then choosing a Template Matching Method.

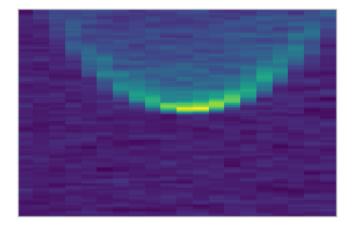
```
spacecraft_id (sample)
```

float32 249.0 249.0 249.0 ... 55.0 55.0

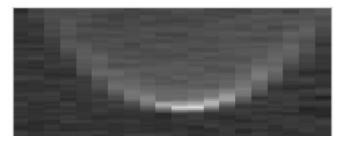
Though it has since been deleted for missing nbrcs/les values, the ddm for sample 1 had a fairly ideal parabolic pattern for a template matching template/ Fortunately, before the sample removal, we took a screenshot of this DDM for use as a template going forward. That file, 'ddm_screenshot.png' is included in the 'Files needed to run' folder.

(56.11)

- 1 #convert to greyscale for better defined parabola
- 2 template_image = cv2.imread(f'{pathTeam}ddm_screenshot.png',0)
- 1 #save a jpg of testing ddm for comparison
- 2 sample selection = 1
- 3 ddm selection = 0
- 4 prepare_test_image(data_set_411, sample_selection, ddm_selection)



- 1 ##again, convert to greyscale
- 2 testing image = cv2.imread('image to test.jpg',0)
- 3 cv2_imshow(testing_image)



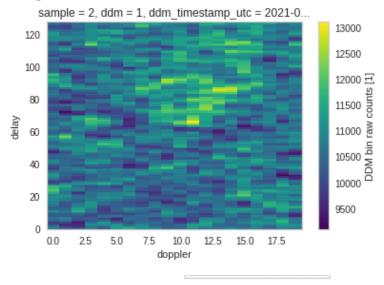
- 1 ##perform template matching
- 2 create_matching(template_image, testing_image)



We looked at DDMs that clearly didn't contain our template image, to show how much lower the maximum template matching coefficient would be:

```
1 ##looking at a ddm much less similar to our template image
2 data_set_411.sel(sample = 2, ddm = 1)['raw_counts'].plot()
```

<matplotlib.collections.QuadMesh at 0x7f86d02b2f50>



```
1 ##make this DDM our testing image:
```

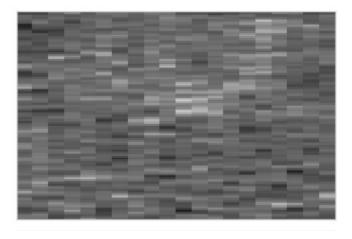
² sample_selection = 2

³ ddm selection = 1

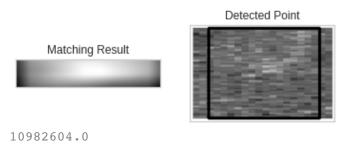
⁴ prepare_test_image(data_set_411, sample_selection, ddm_selection)

⁵ testing_image = cv2.imread('image_to_test.jpg',0)

⁶ cv2 imshow(testing image)



- 1 ##perform template matching
- 2 create_matching(template_image, testing_image)



cv2.TM_CCOEFF_NORMED

Predictably, the maximum template matching coefficient was lower for many of the matching algorithms in this case.



After template matching for all these methods across many DDMs, the team settled on the method 'TM_COEFF_NORMED'. This method's maximum template matching coefficient values seemed to be the most linked with the different kind of visual patterns possible for the delay doppler maps. Once the algorithm was selected, it remained for us to run such a matching, with that method, on every DDM:

```
1 ##Now, create an array of maximum matching coefficients for template matches on the DDM
2 ###NoTE: This cell takes a while to run; Approx 5 minutes for 353 samples
3 coeff_array = match_coeff_array(data_set_411, template_image)

1 #pull this data into a dataframe
2 coeff_frame = pd.DataFrame(coeff_array, columns = ['Max Matching Coeff'])

ATCAACARA 0
1 ##Perform combination of max coefficient values into the original dataset
2 data_set_411 = create_complete_with_maxes(data_set_411, coeff_frame)
3 data_set_411
```

```
▶ Dimensions:
                      (ddm: 4, delay: 128, doppler: 20, sample: 353)
▼ Coordinates:
   sample
                      (sample)
                                                            int64 2 3 4 5 7 ... 1935 1980 1988 2...
   ddm
                      (ddm)
                                                            int64 0 1 2 3
   ddm_timestamp... (sample)
   sp_lat
                      (sample, ddm)
   sp lon
                      (sample, ddm)
▼ Data variables:
   Max Matching ...
                     (sample, ddm)
                                                                                                  DMC rotio
                      (aamala ddm)
```

1 ##delete dummy variables involved in creating the Max Template Matching Coeff Variable: 2 del coeff_array, coeff_frame, template_image, ddm_selection, sample_selection

spacecraft num (sample) float32 2.0 2.0 2.0 2.0 ... 7.0 7.0 7.0

If we search for samples 0 and 1206 (with our junk DDMs), we'll find they were removed automatically by the NBRCS/LES missing value removal function. Therefore, we didn't need to execute any specific code to remove samples with all-zero power values for any DDMs in this case.

wV10m (sample, ddm) tloat64 6.087 7.245 6.24 ... -3.471 -0....

Creating Wind Speed Variable

Now, it remains to convert the wind speed component variables to create one more variable with simple 'Wind Speed', with the Pythagorean Theorem:

```
1 #Calculate/save wind speed from wind vector components
```

- 2 data set 411 = create speed var(data set 411)
- 3 data_set_411

This set is now completely processed, with all relevant DDM calibration/Background Grid data variables, and also, no missing values and no samples associated with junk DDMs. The team performed all these processing steps for 5 days worth of CYGNSS data, across 5 months of sampling in the Full DDM NASA database for 2021. The results were then compiled into a single, large modeling dataset, which can be read in at the start of our next section (which covers modeling).

wU10m (sample, ddm) float64 -2.678 -5.578 ... 10.97 4.847

- Linear Modeling
- Exploratory Analysis on the Modeling Dataset
 - 1 #read in the large modeling dataset
 - 2 ##the file containing this dataset can be found in the 'Files needed to run' folder
 - 3 modeling_set = xr.open_dataset(f'{pathTeam}modeling_dataset.nc')
 - 4 modeling_set

```
▶ Dimensions:
                      (ddm: 4, delay: 128, doppler: 20, sample: 2670)
▼ Coordinates:
                                                            int64 0 1 2 3 4 ... 2666 2667 2668 2...
   sample
                      (sample)
   ddm
                      (ddm)
                                                            int64 0 1 2 3
   ddm_timestamp... (sample)
   sp_lat
                      (sample, ddm)
   sp lon
                      (sample, ddm)
▼ Data variables:
   Max Matching ...
                      (sample, ddm)
```

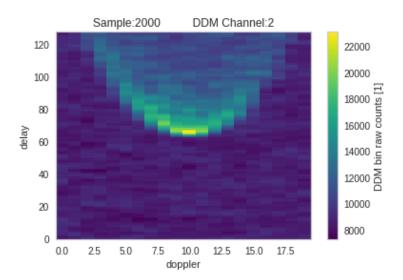
This dataset contains about 10,000 full Delay Doppler Maps across five days of CYGNSS satellite Delay Doppler Mapping sampling intervals, with four Delay Doppler Maps per sample. Furthermore, each day was selected from a myriad of days in a given month, so that 5 months are represented in the dataset: March, April, June, July and August.

```
raw counts (sample ddm delay donnler) float64
```

We began by plotting a Delay Doppler Map (DDM), one of approx. 10,000 in the dataset, to show how each DDM is associated with a particular set of calibration and wind speed/wave height values:

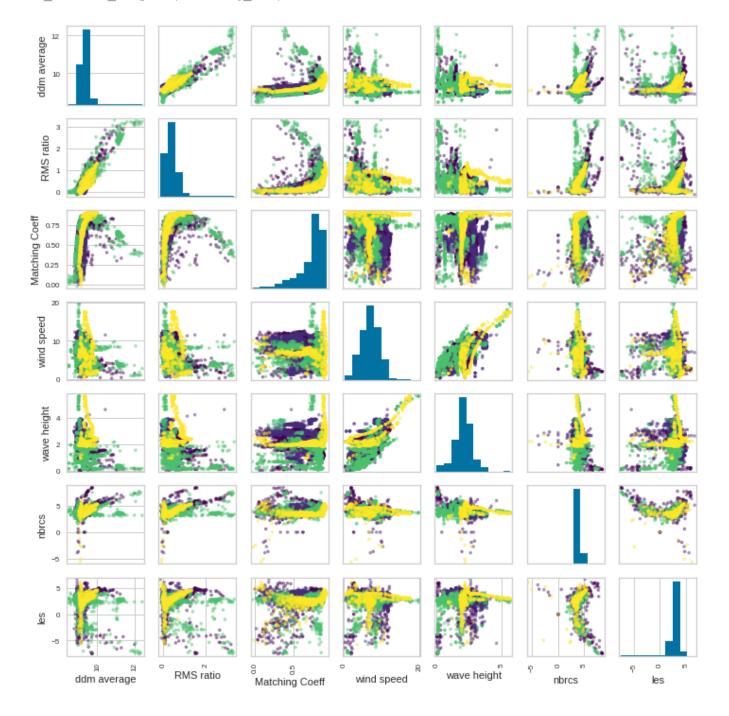
```
1 ##plot ddm with corresponding variable values
2 sample_select = 2000
3 ddm_select = 2
4 ddm_plots_with_vars(modeling_set, sample_select, ddm_select)
5 del sample_select, ddm_select

ddm average: RMS ratio: Max Matching Coeff: wind speed(m/s): wave height:
0 13470.96 1.57 0.867 6.409 1.84
```



The first visualization of interest in our exporatory analysis is a full scatter plot matrix of all these ddm calibration and wind/wave data, to get a sense of any patterns/structure in their pairwise relationships:

- $1\ \#\text{create}$ scatterplot matrix of variables from set, colored by UTC timestamp for the sam
- 2 ###NOTE: les and nbrcs contain negative values and so have been transformed with sign(x
- 3 ###stretch/better visualize their patterns
- 4 ##ALSO: RMS ratio and DDM average values have been transformed with standard log scale
- 5 full_scatter_compare(modeling_set)



A correlation matrix will help to quantify the patterns (or in certain cases, lack thereof) in the plots above:

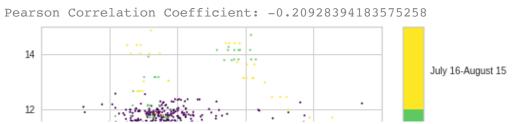
^{1 ##}create a correlation matrix for comparison with the scatterplot matrix given just abo 2 correlation matrix(modeling set)

| | ddm average | RMS ratio | Matching Coeff | wind speed | wave height | nbrcs | les |
|-------------------|----------------|--------------|-------------------|---------------|----------------|--------|--------|
| ddm average | 1.000 | 0.897 | 0.067 | -0.187 | -0.254 | 0.304 | -0.439 |
| RMS ratio | 0.897 | 1.000 | 0.037 | -0.196 | -0.290 | 0.317 | -0.375 |
| Matching Coeff | 0.067 | 0.037 | 1.000 | -0.138 | -0.084 | -0.125 | 0.069 |
| wind speed | -0.187 | -0.196 | -0.138 | 1.000 | 0.679 | -0.209 | 0.006 |
| wave height | -0.254 | -0.290 | -0.084 | 0.679 | 1.000 | -0.198 | 0.080 |

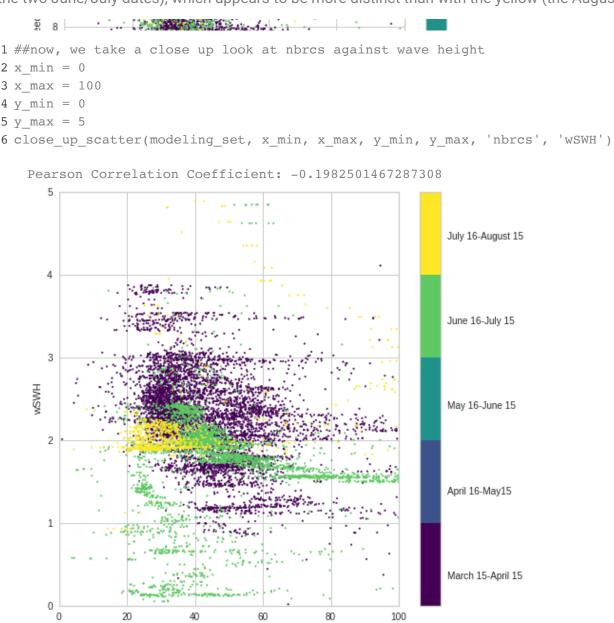
The relatively high correlation between wave height and wind speed is unsurprising. Also, the high correlation between ddm average and RMS ratio is no surprise, as both represent a sort of average of large portions of data in each DDM. That les and nbrcs are related is interesting, although both represent NASA callibrations of the 'raw count' power values that color each DDM.

It seems worthwhile to take a much closer look at the individual scatterplots of 1) nbrcs against wind speed, 2) nbrcs against wave height, 3) RMS ratio against wind speed, 4) RMS ratio against wave height. This is because these pairs of variables represent the most highly correlated pairs of callibration values with wind speed/wave height values:

```
1 #a closer look at nbrcs vs. wind speed
2 x_min = 0
3 x_max = 100
4 y_min = 0
5 y_max = 15
6 close_up_scatter(modeling_set, x_min, x_max, y_min, y_max, 'nbrcs', 'wind speed')
```



The negative relationship seems to be stronger for certain time intervals than others. For example, the purple time interval (the two March-April dates) seems to have a more distinct relationship than the green (the two June/July dates), which appears to be more distinct than with the yellow (the August date).

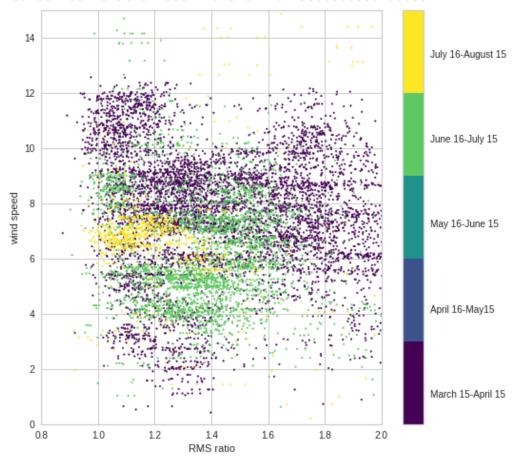


Again, while a slight negative relationship appears overall, that relationship is stronger in the first two time intervals (the spring and summer intervals) than the other two.

nbrcs

```
2 x_min = .8
3 x_max = 2
4 y_min = 0
5 y_max = 15
6 close_up_scatter(modeling_set, x_min, x_max, y_min, y_max, 'RMS ratio', 'wind speed')
```

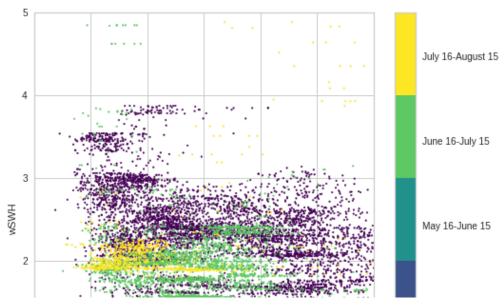
Pearson Correlation Coefficient: -0.19586388337276633



On somewhat closer examination, we see that the negative correlated relationship looks fairly weak, regardless of specific time interval.

```
1 #take a closer look at RMS ratio vs. Wave Height
2 x_min = .8
3 x_max = 2
4 y_min = 0
5 y_max = 5
6 close_up_scatter(modeling_set, x_min, x_max, y_min, y_max, 'RMS ratio', 'wSWH')
```





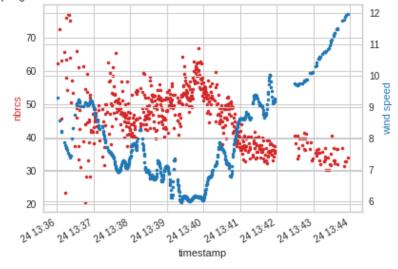
For RMS ratio vs. Wave Height, the inverse correlation appears slightly for especially low or especially high values of RMS ratio, but elsewhere, these variables appear almost independent.

We might be able to clarify some of the patterns by looking at them over specific time intervals. We do that with the following plots, looking at nbrcs/les/RMS ratio against wind speed and wave height in pairwise turns:

RMS ratio

```
1 #create a twin plot with time on mutual x-axis comparing nbrcs and wind speed for sampl
2 #NOTE: this plot isolates trends across just one of this satellite's four DDM channels,
3 start_samp = 0
4 end_samp = 593
5 channel = 0
6 var_of_intA = 'nbrcs'
7 var_of_intB = 'wind speed'
8 create_twin_plot(modeling_set, start_samp, end_samp, channel, var_of_intA, var_of_intB)
9 del start samp, end samp, channel, var of intA, var of intB
```

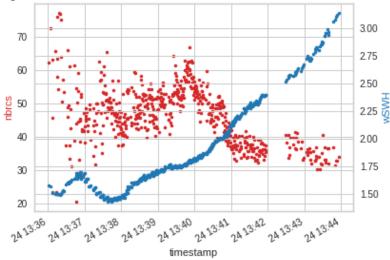
Sampling Interval:2021-03-24T13:36:01.999261583 to 2021-03-24T13:43:55.999261681



In just this isolated time interval, the anti-correlation of the nbrcs and wind speed variables seems fairly strong. Let's look at nbrcs vs. wave height, for the same time interval and DDM channel:

```
1 start_samp = 0
2 end_samp = 593
3 channel = 0
4 var_of_intA = 'nbrcs'
5 var_of_intB = 'wSWH'
6 create_twin_plot(modeling_set, start_samp, end_samp, channel, var_of_intA, var_of_intB)
7 del start_samp, end_samp, channel, var_of_intA, var_of_intB
```

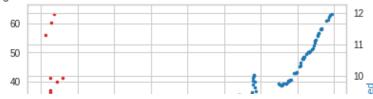
Sampling Interval:2021-03-24T13:36:01.999261583 to 2021-03-24T13:43:55.999261681



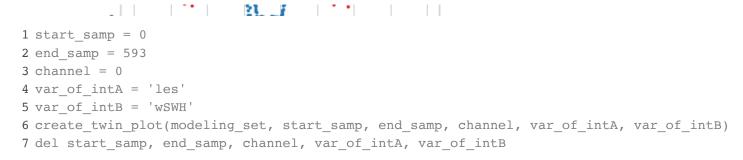
Here, while the significant wave height plot layout is a much denser than wind speed, the anti-correlation is still more obvious than when we plot all ddm data across all time intervals, as with the scatter plots above. Now, we take a look at 'les' vs. wind speed and wave height.

```
1 start_samp = 0
2 end_samp = 593
3 channel = 0
4 var_of_intA = 'les'
5 var_of_intB = 'wind speed'
6 create_twin_plot(modeling_set, start_samp, end_samp, channel, var_of_intA, var_of_intB)
7 del start_samp, end_samp, channel, var_of_intA, var_of_intB
```

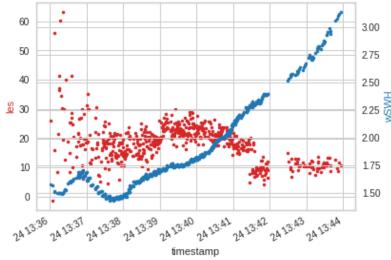
Sampling Interval:2021-03-24T13:36:01.999261583 to 2021-03-24T13:43:55.999261681



The anticorrelation here is still present but a bit less strong than for nbrcs, which is mimicked by the anticorrelation of these variables over all the data in the modeling dataset. Now, looking at les vs. wave height:



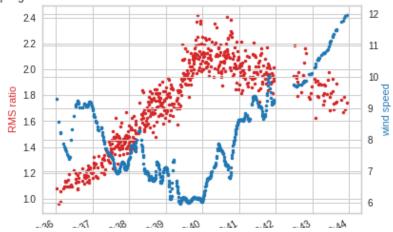
Sampling Interval:2021-03-24T13:36:01.999261583 to 2021-03-24T13:43:55.999261681



For these variables, the relationship is far more slight and appears to almost reverse halfway through the time interval. Finally, we take examine the relationship between RMS ratio (our most promising DDM calibration done by the team), and wind speed/wave height over these intervals:

```
1 start_samp = 0
2 end_samp = 593
3 channel = 0
4 var_of_intA = 'RMS ratio'
5 var_of_intB = 'wind speed'
6 create_twin_plot(modeling_set, start_samp, end_samp, channel, var_of_intA, var_of_intB)
7 del start samp, end samp, channel, var of intA, var of intB
```

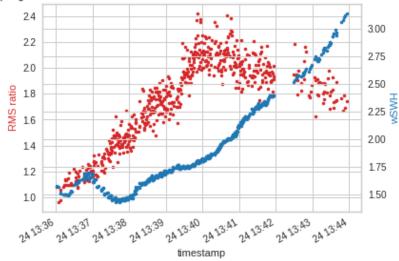
Sampling Interval:2021-03-24T13:36:01.999261583 to 2021-03-24T13:43:55.999261681



Once more, there is a more well-defined anti-correlation between RMS ratio and wind speed over this smaller time interval than over the whole dataset. Looking at RMS ratio vs. wave height, we get:

```
1 start_samp = 0
2 end_samp = 593
3 channel = 0
4 var_of_intA = 'RMS ratio'
5 var_of_intB = 'wSWH'
6 create_twin_plot(modeling_set, start_samp, end_samp, channel, var_of_intB)
7 del start_samp, end_samp, channel, var_of_intA, var_of_intB
```

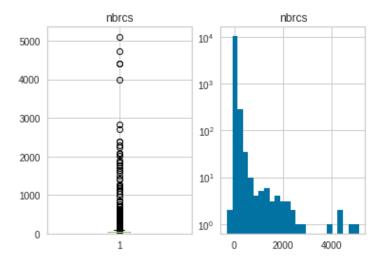
Sampling Interval:2021-03-24T13:36:01.999261583 to 2021-03-24T13:43:55.999261681



As with les, we get a less obvious relationship with wave height than with wind speed.

At this point, it may be helpful to look at the distributions of these individual variables as well, along with basic statistical summaries for each variable. NOTE: these represent data across the entire dataset, rather than across a limited time interval.

1 #create the boxplot and corresponding histogram for nbrcs
2 #note, the histogram frequency axis is converted to a log scale for visual clarity
3 ddm box plot(modeling set, 'nbrcs')

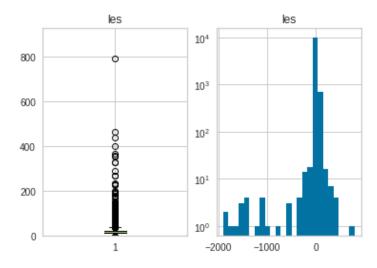


1 stat_summaries(modeling_set, 'nbrcs')

| | Statistics | Values |
|---|------------|----------|
| 0 | Mean | 59.007 |
| 1 | Median | 42.214 |
| 2 | Std | 139.520 |
| 3 | Max | 5092.975 |
| 4 | Min | -289.579 |

It seems that nbrcs has a very large number of outliers, with most data being concentrated around a very few values between zero and 1. Once we transform the frequency axis with a log scale, we see that overall, nbrcs is non-normally distributed, being fairly right-skewed.

- 1 #create the boxplot and corresponding histogram for les
- 2 ##As with nbrcs, the distribution for les seems to be heavily concentrated around zero,
- 3 ddm_box_plot(modeling_set, 'les')



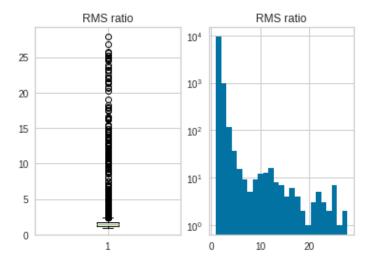
1 stat_summaries(modeling_set, 'les')

Statistics Values

```
0 Mean 15.991
1 Median 15.458
2 Std 64.356
3 Max 791.867
4 Min -1904.772
```

The distribution for les seems to be heavily concentrated around zero, so that the histogram also benefits from a log scale transformation, with the more visible distribution on the right showing a somewhat left-skewed graphic.

- 1 #create the boxplot and corresponding histogram for RMS ratio
- 2 ##As with nbrcs and les, RMS ratio has been given a log scale density axis for its hist
- 3 ddm_box_plot(modeling_set, 'RMS ratio')

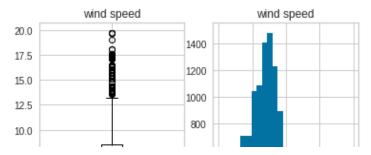


1 stat_summaries(modeling_set, 'RMS ratio')

| | Statistics | Values |
|---|------------|--------|
| 0 | Mean | 1.644 |
| 1 | Median | 1.400 |
| 2 | Std | 1.526 |
| 3 | Max | 27.831 |
| 4 | Min | 0.875 |

Here, we see that RMS ratio is extremely right-skewed, with mean greater than the median. Finally, we look at the individual distributions and statistical summaries of wind speed and wave height:

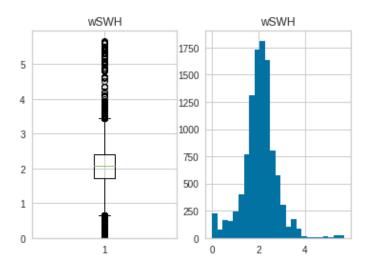
- 1 #create box plot and corresponding histogram for wind speed:
- 2 ddm_box_plot(modeling_set, 'wind speed')



1 stat_summaries(modeling_set, 'wind speed')

| | Statistics | Values |
|---|------------|--------|
| 0 | Mean | 6.988 |
| 1 | Median | 7.078 |
| 2 | Std | 2.474 |
| 3 | Max | 19.721 |
| 4 | Min | 0.167 |

- 1 ##and for wave height...
- 2 ddm_box_plot(modeling_set, 'wSWH')



1 stat_summaries(modeling_set, 'wSWH')

| | Statistics | Values |
|---|------------|--------|
| 0 | Mean | 2.066 |
| 1 | Median | 2.077 |
| 2 | Std | 0.704 |
| 3 | Max | 5.674 |
| 4 | Min | 0.017 |

Both these variables are fairly normal in distribution. The presence of some correlations and anticorrelations in the data, particularly across specific time intervals, but also in the dataset at large, would suggest that linear modeling might be fruitful, given the right combination of non-colinear variables.

Next, we begin the process of assessing all DDM callibration and wind/wave variables in the modeling

Assessing Colinearity

- 1 ##start by creating a pandas dataframe from the ddm calibrations/wind speed/wave height
- 2 modeling_frame = create_dataframe(modeling_set)
- 3 modeling_frame = modeling_frame.astype('float64')
- 4 modeling_frame = modeling_frame.round(3)
- 5 modeling_frame

| | ddm average | RMS ratio | Matching Coeff | wind speed | wave height | nbrcs | les |
|-------|-------------|-----------|----------------|------------|-------------|--------|--------|
| 0 | 7401.46 | 1.077 | 0.289 | 9.300 | 1.577 | 62.245 | 25.984 |
| 1 | 14560.06 | 1.648 | 0.757 | 8.765 | 1.721 | 33.157 | 12.925 |
| 2 | 10968.92 | 1.927 | 0.860 | 9.553 | 1.247 | 38.311 | 20.428 |
| 3 | 15699.80 | 1.839 | 0.845 | 6.573 | 1.903 | 31.601 | 13.440 |
| 4 | 7512.64 | 0.964 | 0.294 | 8.576 | 1.573 | 72.595 | -1.148 |
| | | | | | | | |
| 10675 | 8693.80 | 1.090 | 0.480 | 7.692 | 2.011 | 60.153 | 36.688 |
| 10676 | 19592.18 | 2.520 | 0.860 | 8.318 | 1.973 | 75.276 | 36.031 |
| 10677 | 9361.48 | 1.614 | 0.853 | 1.935 | 2.103 | 88.845 | 42.700 |
| 10678 | 8453.16 | 1.104 | 0.229 | 6.609 | 1.954 | 0.000 | 0.000 |
| 10679 | 8374.92 | 1.045 | 0.347 | 7.691 | 2.011 | 49.668 | 22.902 |

¹⁰⁶⁸⁰ rows × 7 columns

1 #remove the dependent variable of interest, in this case, 'wind speed', fit multiple re 2 find_VIF(modeling_frame, 'wind speed')

| | feature | VIF |
|---|----------------|--------|
| 0 | ddm average | 16.517 |
| 1 | RMS ratio | 11.590 |
| 2 | Matching Coeff | 9.568 |
| 3 | wave height | 6.484 |
| 4 | nbrcs | 1.974 |
| 5 | les | 1.988 |

It actually does seem that given the general rule of of thumb that a VIF value should not exceed 10 (such a high value indicates a colinearity problem for that variable), we do have possible collinearity issues. We

resolve this by removing ddm average and re-runing the multiple regression, re-checking the Variance Inflation Factors again, afterwards:

```
1 #drop the ddm average variable and reassess VIF:
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = remove_colinear_var(modeling_frame, 'ddm average')
4 find VIF(modeling frame, 'wind speed')
```

| | feature | VIF |
|---|----------------|-------|
| 0 | RMS ratio | 2.506 |
| 1 | Matching Coeff | 8.651 |
| 2 | wave height | 6.393 |
| 3 | nbrcs | 1.961 |
| 4 | les | 1.861 |

The removal of ddm average appears to have corrected the colinearity problem (all VIF values are now under 10), suggesting that it was the extremely high correlation of RMS ratio to ddm average that was making our original VIF assessment model so highly multicolinear (ddm average and RMS ratio would have explained roughly the same amount of variance in our eventual wind speed model).

We repeated the VIF assessment process for a potential model with wave height as the dependent variable:

```
1 ##reset modeling dataframe, but this time find VIF with all variables except Wave Heigh
2 modeling_frame = create_dataframe(modeling_set)
3 find_VIF(modeling_frame, 'wave height')
```

```
        feature
        VIF

        0
        ddm average
        16.380

        1
        RMS ratio
        11.401

        2
        Matching Coeff
        9.056

        3
        wind speed
        5.837

        4
        nbrcs
        1.979

        5
        les
        1.983
```

```
1\ \mbox{\#drop} the ddm average variable and reassess VIF:
```

² modeling frame = create dataframe(modeling set)

```
3 modeling_frame = remove_colinear_var(modeling_frame, 'ddm average')
4 find_VIF(modeling_frame, 'wave height')
```

| | feature | VIF |
|---|----------------|-------|
| 0 | RMS ratio | 2.484 |
| 1 | Matching Coeff | 7.916 |
| 2 | wind speed | 5.803 |
| 3 | nbrcs | 1.967 |
| 4 | امم | 1 863 |

And once again, we see the colinearity problem corrected. So we see that in any linear model built to predict either wind speed or wave height, we would certainly want to exclude the crude ddm average, as it is too highly correlated with RMS ratio, and because, of the two variables, RMS ratio offers more promising correlation with our would-be dependent variables.

If we wanted to be even more aggressive in our insistence that the model have no multicolinearity problems (some statisticians insist on having VIFs not much higher than 5), we could further remove Maximum Template Matching Coefficient as a variable from these VIF calculations and see the result:

```
1 #remove ddm average and max matching coeff from modeling dataframe and assess VIF for w
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = remove_colinear_var(modeling_frame, 'ddm average')
4 modeling_frame = remove_colinear_var(modeling_frame, 'Matching Coeff')
5 find VIF(modeling_frame, 'wind_speed')
```

| | feature | VIF |
|---|-------------|-------|
| 0 | RMS ratio | 1.977 |
| 1 | wave height | 2.088 |
| 2 | nbrcs | 1.909 |
| 3 | les | 1.723 |

```
1 #do the same for a model that would have wave height as the dependent variable
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = remove_colinear_var(modeling_frame, 'ddm average')
4 modeling_frame = remove_colinear_var(modeling_frame, 'Matching Coeff')
5 find VIF(modeling frame, 'wave height')
```

For both our prospect models (using wind speed or wave height as the dependent variable), we would drastically mitigate the colinearity problems we would otherwise get by removing both ddm average and maximum template matching coefficient as independent regressors to those models, though whether we should remove Maximum matching coefficient too depends on how strongly we wish to avoid the problem of multicolinearity.

For the purposes of our model, we won't remove Maximum Matching Coefficient, as its colinearity with wind speed and wave height is far more limited than when we include ddm average.

Variable Selection

Now that we have identified a set of variables that is not colinear for linear models that might use wind speed or wave height as a predictand, we can do best subsets variable selection to try and identify a 'best' combination of regressor/predictor variables for such models.

```
1 ##start by building a dataframe with only the variables which we know are not colinear
2 #also in this cell, we define y (the dependent variable for the model), as wind speed
3 modeling frame = create dataframe(modeling set)
4 modeling frame = modeling frame.drop(['ddm average'], axis = 1)
5 y = modeling frame['wind speed']
6 X = modeling frame.drop(['wind speed'], axis = 1)
1 ##use functions defined above to create a dataframe that contains best models produced
2 models_highest_RSS = pd.DataFrame(columns=["RSS", "model"])
3 for i in range(1,6):
     models_highest_RSS.loc[i] = highest_RSS(i)
5 models highest RSS vals = models highest RSS['RSS']
6 models highest RSS vals = models highest RSS vals.astype('float64')
7 models highest RSS vals.round(3)
8 ##This cell produces a dataframe with the number of variables in the model on the left
9 ###with that number of variables
        39928.029
   2
       37834.713
   3
       37492.034
        37279.259
       36970.417
   Name: RSS, dtype: float64
1 ##get more information about the models that produced the highest RSS for each number o
2 ###start with model that produced highest RSS for model with only 1 regressor:
```

OLS Regression Results

3 print(models highest RSS.loc[1, "model"].summary())

```
Dep. Variable: wind speed R-squared (uncentered):
                                                   0.9
                     OLS Adj. R-squared (uncentered):
Model:
                                                   0.9
           Least Squares F-statistic:
                                                 1.463e+
Method:
            Tue, 21 Dec 2021 Prob (F-statistic):
                                                    0.
Date:
                 01:33:53 Log-Likelihood:
                                                  -2219
Time:
No. Observations:
                   10680 AIC:
                                                 4.439e+
Df Residuals:
                   10679 BIC:
                                                 4.440e+
Df Model:
                      1
Covariance Type: nonrobust
______
          coef
               std err t P>|t| [0.025 0.975]
______
wave height 3.2786 0.009 382.508 0.000
                                      3.262
______
                  560.889 Durbin-Watson:
Omnibus:
                                             1.071
Prob(Omnibus):
                   0.000 Jarque-Bera (JB):
                                            761.290
                   -0.498 Prob(JB):
                                           4.88e-166
Skew:
                                            1.00
Kurtosis:
                   3.847 Cond. No.
```

Warnings:

Dep. Variable:

Model:

Method:

[1] Standard Errors assume that the covariance matrix of the errors is correctly spec

It is not too surprising that any model including wave height, even one with just the variable wave height, (which is so highly correlated with wind speed), will have a very high R^2 statistic. It will be interesting, then, to do the same best subsets selection process, but on all non-colinear variables not including wave height:

```
1 ##start by building a dataframe with only the variables which we know are not colinear
2 ###This time, we also exclude wave height as a predictor:
3 #also in this cell, we define y (the dependent variable for the model), as wind speed
4 modeling frame = create dataframe(modeling set)
5 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
6 y = modeling frame['wind speed']
7 X = modeling frame.drop(['wind speed'], axis = 1)
8 X = X.drop(['wave height'], axis = 1)
1 ##use functions defined above to create a dataframe that contains best models produced
2 models highest RSS = pd.DataFrame(columns=["RSS", "model"])
3 for i in range(1,5):
     models highest RSS.loc[i] = highest RSS(i)
1 ##again, we want more specific information about the model that uses only 1 regressor v
2 ##for the models with 3 variables:
3 print(models_highest_RSS.loc[1, "model"].summary())
                                    OLS Regression Results
```

Least Squares F-statistic:

wind speed R-squared (uncentered):

OLS Adj. R-squared (uncentered):

0.8

0.8

5.103e+

| Date: | Tue, 2 | 1 Dec 2021 | <pre>Prob (F-statistic):</pre> | | | 0. |
|---|--------|-------------------|--------------------------------|----------|----------|---------|
| Time: | | 01:33:54 | Log-Likelihood: | | | -2718 |
| No. Observations: | | 10680 | AIC: | | | 5.437e+ |
| Df Residuals: | | 10679 | BIC: | | | 5.437e+ |
| Df Model: | | 1 | | | | |
| Covariance Type: | | nonrobust | | | | |
| ======================================= | | std err | t | P> t | [0.025 | 0.975] |
| Matching Coeff | 9.1541 | 0.041 | 225.897 | 0.000 | 9.075 | 9.234 |
| Omnibus: | | 504.611 | Durbin-Wat | son: | | 1.588 |
| Prob(Omnibus): 0.000 | | Jarque-Bera (JB): | | 578.962 | | |
| Skew: | | 0.551 | <pre>Prob(JB):</pre> | | 1.91 | e-126 |
| Kurtosis: | | 3.293 | Cond. No. | | | 1.00 |
| =========== | | | ======== | ======== | ======== | ===== |

Warnings:

[1] Standard Errors assume that the covariance matrix of the errors is correctly spec

1 ##and the best model with 2 variables

2 print(models_highest_RSS.loc[2, "model"].summary())

OLS Regression Results

| Dep. Variable: Model: Method: Date: Time: No. Observations: Df Residuals: Df Model: | Lea | OLS st Squares | R-squared (uncentered): Adj. R-squared (uncentered): F-statistic: Prob (F-statistic): Log-Likelihood: AIC: BIC: | | | 0.8 0.8 2.559e+ 0. -2716 5.434e+ 5.436e+ |
|---|--------|------------------------------------|---|-------|---------------------------------------|--|
| Covariance Type: | | nonrobust | | | | |
| | coef | std err | t | P> t | [0.025 | 0.975] |
| 3 | 9.2343 | | 213.327 -5.223 | 0.000 | 9.149 | |
| Omnibus: Prob(Omnibus): Skew: Kurtosis: | | 532.527 0.000 0.566 3.318 | Durbin-Watson: Jarque-Bera (JB): Prob(JB): Cond. No. | | 1.589 616.027 1.70e-134 220. | |

Warnings:

[1] Standard Errors assume that the covariance matrix of the errors is correctly spec

1 #and the model with 3 regressors:

2 print(models_highest_RSS.loc[3, "model"].summary())

OLS Regression Results

wind speed R-squared (uncentered):

0.8

| Model: | OLS Adj. R-squared (uncentered): | | | | | 0.8 |
|----------------------|--------------------------------------|------------|----------------|--|----------------------|-----------------------|
| Method: | Lea | st Squares | | 1.708e+ | | |
| Date: | Tue, 21 Dec 2021 Prob (F-statistic): | | | | | 0. |
| Time: | | 01:33:54 | Log-Likeli | -2716 | | |
| No. Observations: | | 10680 | AIC: | 5.433e+ | | |
| Df Residuals: | | 10677 | BIC: | | | 5.435e+ |
| Df Model: | | 3 | | | | |
| Covariance Type: | | nonrobust | | | | |
| | | | | | ======== | |
| | coef | std err | t | P> t | [0.025 | 0.975] |
| Matching Coeff | 9.3291 | 0.050 | 184.756 | 0.000 | 9.230 | 9.428 |
| _ | -0.0017 | 0.000 | -6.352 | 0.000 | -0.002 | -0.001 |
| les | -0.0022 | 0.001 | -3.643 | 0.000 | -0.003 | -0.001 |
| Omnibus: | | 567.718 | Durbin Wat | ====================================== | :======= | 1.590 |
| | | 0.000 | Durbin-Watson: | | | |
| Prob(Omnibus): Skew: | | | Prob(JB): | | 663.290 9.30e-145 | |
| Kurtosis: | 0.586 3.343 | | Cond. No. | | 263. | |
| | | 3.343 | ========= | .======== | .======== | 203 . ===== |

Warnings:

[1] Standard Errors assume that the covariance matrix of the errors is correctly spec

- 1 #the model that regresses wind speed on all four variables
- 2 print(models_highest_RSS.loc[4, "model"].summary())

OLS Regression Results

| Dep. Variable: | | wind speed | R-squared | 0.8 | | | |
|---|----------------------------------|------------|---------------------|----------|-----------|-----------|--|
| Model: | OLS Adj. R-squared (uncentered): | | | | | 0.8 | |
| Method: | Lea | st Squares | F-statisti | 1.282e+ | | | |
| Date: | Tue, 2 | 1 Dec 2021 | Prob (F-statistic): | | | 0. | |
| Time: | | 01:33:54 | Log-Likeli | -2716 | | | |
| No. Observations: | 10680 AIC: | | | | 5.433e+ | | |
| Df Residuals: | | 10676 | BIC: | 5.436e+ | | | |
| Df Model: | | 4 | | | | | |
| Covariance Type: | | nonrobust | | | | | |
| | | | | | | | |
| | | | | | [0.025 | | |
| RMS ratio | | | | | -0.087 | | |
| Matching Coeff | 9.4289 | 0.068 | 138.706 | 0.000 | 9.296 | 9.562 | |
| nbrcs | -0.0016 | 0.000 | -5.913 | 0.000 | -0.002 | -0.001 | |
| les | -0.0025 | 0.001 | -4.018 | 0.000 | -0.004 | -0.001 | |
| Omnibus: | ======= | 582.012 | Durbin-Watson: | | 1.589 | | |
| Prob(Omnibus): | | 0.000 | Jarque-Bera (JB): | | 682.783 | | |
| Skew: | | 0.593 | Prob(JB): | | 5.44 | 5.44e-149 | |
| Kurtosis: | | 3.354 | Cond. No. | | | 363. | |
| ======================================= | ======= | ======== | | :======= | | ===== | |

Warnings:

[1] Standard Errors assume that the covariance matrix of the errors is correctly spec

Clearly, the exclusion of wave height reduces the adjusted R^2 value for the model, but it appears that our best model with 3 variables (from the standpoint of maximizing RSS), regresses wind speed on the variables les, nbrcs, and Matching coeff. Were we to regress wind speed on just 2 variables, we should use nbrcs and Matching coeff. Were we to use only 1 variable, we would see that Matching coeff is the best variable to regess wind speed on. We also see that using all four variables maximizes adjusted R^2 .

Next, we do the same best subsets selection to maximize RSS for models predicting wave height:

```
1 ##fit models for best subsets variable selection for wave height prediction, dropping w
 2 modeling frame = create dataframe(modeling set)
 3 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
 4 y = modeling_frame['wave height']
 5 X = modeling frame.drop(['wave height'], axis = 1)
 6 X = X.drop(['wind speed'], axis = 1)
7 models highest RSS = pd.DataFrame(columns=["RSS", "model"])
8 for i in range(1,5):
     models_highest_RSS.loc[i] = highest_RSS(i)
10 #find the 'best' model that regresses wind speed on each number of variables
11 #save the R^2 value:
12 rsquared = []
13 for i in range(1,5):
14 rsquared.append(models highest RSS.loc[i, "model"].rsquared)
15 rsquared
    [0.842, 0.843, 0.844, 0.844]
```

It seems that in both cases, where we are trying to predict wind speed or wave height, of the models which maximize RSS for each number of regressor variables, the model that regresses on all four possible predictors is the one that maximizes adjusted \mathbb{R}^2 .

```
1 del models_highest_RSS, models_highest_RSS_vals, rsquared
```

Checking Error Assumptions and Identifying Influential Observations/Outliers

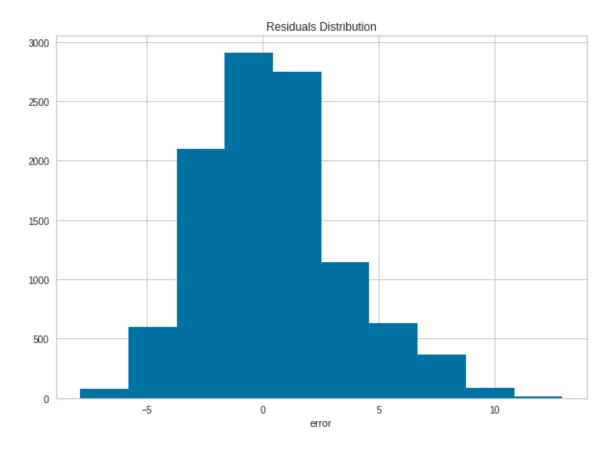
In this section, we must check that the error assumptions in Linear Regression Analysis hold for our data. First, it must be true that errors/residuals (differences between the actual values and the predicted ones) are roughly normally distributed. Second, it must be the case that the errors are basically uncorrelated. Lastly, it must hold true that the variances of the errors are more-or-less constant (homoscedasticity).

Checking Error Assumptions for the Wind Speed Model:

```
1 #first we build our modeling dataframe that includes all the four variables we want to
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
```

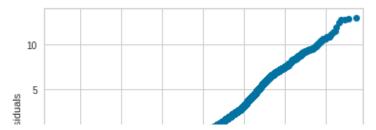
```
4 modeling_frame = modeling_frame.drop(['wave height'], axis = 1)
5 y = modeling_frame['wind speed']
6 X = modeling_frame.drop(['wind speed'], axis = 1)
7 #fit our model for wind speed/summarize:
8 mod = sm.OLS(y,X)
9 regress = mod.fit()
10 mod resid = regress.resid
```

1 #get the residuals and create a histogram of their distribution:
2 resid Hist(regress)



This histogram of error values from our model seems to confirm that the assumption about rough normality of errors distribution is satisfied. We can double-check this assumption with the use of a QQ plot:

```
1 ##create a QQ plot to assess normality of errors/residuals from the model
2 sm.qqplot(mod_resid, ylabel = 'residuals')
3 py.show()
```



Here, we see that the qq plot shows a fairly straight line though not perfect, which indicates a fairly normal distribution; (curvalinear structure in the qq plot would indicate non-normal residuals). So the first of the error assumptions, that of basic normality would at first appear to indicate our first error assumption is satisfied.

rneoretical Quantiles

Next, we look at the assumption of our errors being uncorrelated, which assess by looking at the Durbin-Watson test statistic in the model summary. According to TowardsDataScience.com page "Verifying the Assumptions of Linear Regression", if the Durbin-Watson test statistic is < 2, there is significant positive residual auto-correlation; if > 2, then there is significant negative autocorrelation; if roughly equal to 2, then there is no autocorrelation and our assumption is satisfied:

```
1 #look at Durbin Watson Statistic from our model summary:
2 regress.summary()
```

OLS Regression Results

Dep. Variable: wind speed R-squared (uncentered): 0.828 Model: OLS Adj. R-squared (uncentered): 0.828 Method: Least Squares F-statistic: 1.282e+04 Date: Tue, 21 Dec 2021 Prob (F-statistic): 0.00 Time: 01:33:54 Log-Likelihood: -27160. No. Observations: 10680 AIC: 5.433e+04 **Df Residuals:** 10676 BIC: 5.436e+04

Df Model: 4

Covariance Type: nonrobust

 coef
 std err
 t
 P>ItI
 [0.025 0.975]

 RMS ratio
 -0.0459 0.021 -2.192
 0.028 -0.087 -0.005

 Matching Coeff
 9.4289 0.068 138.706 0.000 9.296 9.562

 nbrcs
 -0.0016 0.000 -5.913 0.000 -0.002 -0.001

 les
 -0.0025 0.001 -4.018 0.000 -0.004 -0.001

 Omnibus:
 582.012
 Durbin-Watson:
 1.589

 Prob(Omnibus):
 0.000
 Jarque-Bera (JB):
 682.783

 Skew:
 0.593
 Prob(JB):
 5.44e-149

 Kurtosis:
 3.354
 Cond. No.
 363.

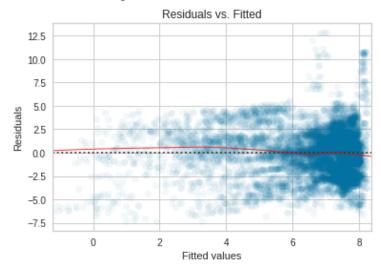
Warnings:

[1] Standard Errors assume that the covariance matrix of the errors is correctly specified.

The Durbin-Watson test statistic indicates that we have some positive autocorrelation between the residuals themselves. While a Durbin-Watson test value less than 1 would be cause for much more concern (according to TowardsDataScience.com), the assumption of uncorrelated errors for our model is not entirely satisfied <u>Source</u>.

Finally, we looked at the last assumption, that the variance of the errors is basically constant, which we do with a scatter plot of residiuals vs. fitted values in the model:

- 1 ##plot a residuals vs. fitted values plot for the model: 2 FitvResid(regress, X, y)
 - /usr/local/lib/python3.7/dist-packages/seaborn/_decorators.py:43: FutureWarning: Pass FutureWarning



A residuals vs. fitted values plot is used to check for constant error variance and for model structure in Linear Regression Modeling. It is a plot of fitted values (predicted \hat{y}) against the residuals, the differences between each observed dependent variable value and the \hat{y} predicted value lying along the line of fit. Ideally, there should be a basically random distribution of points around the mean for the residuals (always zero in a proper linear regression model). Source: Faraway, *Linear Models with R*, 2015, pg 77.

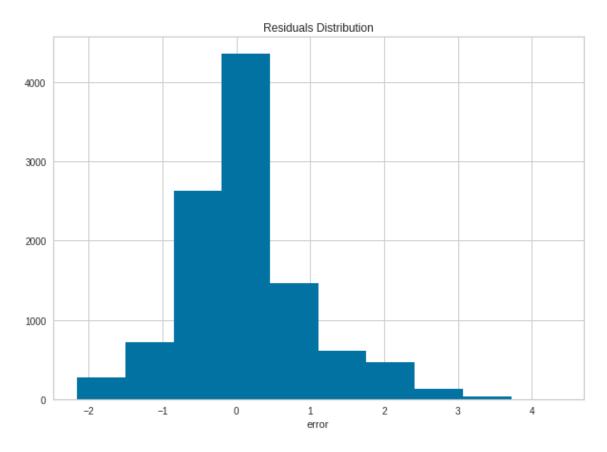
The variance of the errors does appear to be roughly constant, despite a bit pattern/structure in the densest regions of this plot. The loess smoother line approximates the horizontal line at Residual = 0 (representing our line of fit in the model). So the error mean does appear to be zero and the constant variance error assumption appears to be mostly satisfied.

Checking Error Assumptions for the Wave Height Model:

1 #first we build our modeling dataframe that includes all the four variables we want to

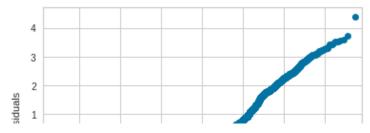
```
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
4 modeling_frame = modeling_frame.drop(['wind speed'], axis = 1)
5 y = modeling_frame['wave height']
6 X = modeling_frame.drop(['wave height'], axis = 1)
7 mod = sm.OLS(y,X)
8 regress = mod.fit()
9 mod_resid = regress.resid

1 #get the residuals and create a histogram of their distribution:
2 resid_Hist(regress)
```



The histogram for residuals in the wave height model seems to be a bit skewed, more so than for the residuals in the wind speed model. Having a bit more curved distribution of points in the qqplot for residuals confirms this:

```
1 ##create a QQ plot to assess normality of errors/residuals from the model
2 sm.qqplot(mod_resid, ylabel = 'residuals')
3 py.show()
```



Now we look at the Durbin-Watson Statistic to assess this model's assumption of non-correlated consecutive errors:

1 ##Now we look at the Durbin-Watson Statistic

OLS Regression Results

Dep. Variable: wave height R-squared (uncentered): 0.844 OLS Model: Adj. R-squared (uncentered): 0.844 Least Squares F-statistic: Method: 1.439e+04 Date: Tue, 21 Dec 2021 Prob (F-statistic): 0.00 Time: 01:34:05 Log-Likelihood: -13585. No. Observations: 10680 AIC: 2.718e+04 **Df Residuals:** 10676 BIC: 2.721e+04

Df Model: 4

2 regress.summary()

4

Covariance Type: nonrobust

 coef
 std err
 t
 P>ItI
 [0.025
 0.975]

 RMS ratio
 -0.0584
 0.006
 -9.953
 0.000
 -0.070
 -0.047

 Matching Coeff
 2.8297
 0.019
 148.381
 0.000
 2.792
 2.867

 nbrcs
 0.0001
 7.72e-05
 1.315
 0.189
 -4.98e-05
 0.000

 les
 0.0005
 0.000
 2.642
 0.008
 0.000
 0.001

 Omnibus:
 910.356
 Durbin-Watson:
 1.262

 Prob(Omnibus):
 0.000
 Jarque-Bera (JB):
 1511.850

 Skew:
 0.632
 Prob(JB):
 0.00

 Kurtosis:
 4.342
 Cond. No.
 363.

Warnings:

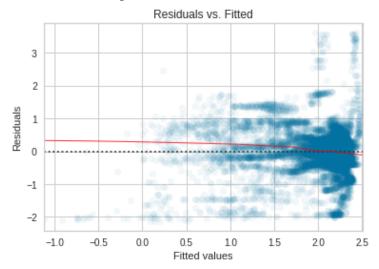
[1] Standard Errors assume that the covariance matrix of the errors is correctly specified.

The Durbin-Watson test statistic is a bit lower than would be ideal, but still above 1. Were it beneath 1, we would have more serious cause to believe that our model's violation of linear modeling error assumptions is deal-breaking.

Finally, we look at a plot of residuals vs. fitted values, to assess the assumption of constant variance for the residuals:

```
1 #get fitted values to plot and plot against residuals
2 FitvResid(regress, X, y)
```

/usr/local/lib/python3.7/dist-packages/seaborn/_decorators.py:43: FutureWarning: Pass FutureWarning



The deviation of the lowess smoother line from the error mean line at Residuals = 0 suggest that the assumption of constant error variance seems more generally violated than we saw with the model predicting wind speed.

Ultimately, there are arguable violations of error assumptions in our models. For the wind speed model, these violations are not as extreme as for the wave height model. From here, we move on to the process of identifying outliers/Influential values.

```
1 del mod resid
```

Identifying Outliers/Influential Values

The Wind Speed Model

```
1 #create the modeling set and fit a model regressing wind speed on variables RMS ratio,
2 #first we build our modeling dataframe that includes all the four variables we want to
3 modeling_frame = create_dataframe(modeling_set)
4 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
5 modeling_frame = modeling_frame.drop(['wave height'], axis = 1)
6 y = modeling_frame['wind speed']
7 X = modeling_frame.drop(['wind speed'], axis = 1)
8 #fit our model for wind speed/summarize:
9 mod = sm.OLS(y,X)
10 regress = mod.fit()
```

First, we will want to look at outlier observations without accounting for their leverage values status. We can do this with a Bonferroni outlier test, which our model has a built-in function to perform:

```
1 #conduct Bonferroni Outlier Test for our wind speed Model:
2 ##NOTE: This cell requires a few minutes to run
3 bonf test = regress.outlier test()
1 #determine which observation has the highest studentized residual:
2 bonf outliers = bonf outlier(bonf test)
3 print(bonf outliers)
         student resid unadj p bonf(p)
   0
               2.206 0.027
                                1.0
                1.939 0.053
                                   1.0
   8
                1.747 0.081
                                  1.0
                1.960 0.050
  12
                                  1.0
                1.751 0.080
  16
                                  1.0
                 . . .
                                   . . .
               1.523 0.128
  10671
                                  1.0
                1.046 0.295
  10674
                                   1.0
  10675
                1.106
                        0.269
                                  1.0
  10678
                1.463 0.143
                                  1.0
  10679
                1.495
                        0.135
                                   1.0
  [1765 rows x 3 columns]
1 #determine ratio of number of outliers in our model to the number of observations, base
2 observation number = 10680
3 outlier_ratio = len(bonf_outliers)/observation_number
4 print(format(outlier ratio, '.3f'))
  0.165
```

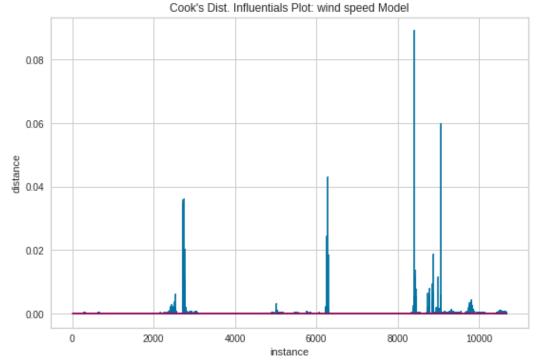
Clearly, we have a rather large number of outliers (about 16.5% of our dataset) by the Bonferroni test criterion. We must now assess whether that translates to a large number of influential observations (which have significant bearing on model fit).

Influential observations are observations whose exclusion would significantly alter the model fit; they account for not just outlier status, but also leverage value status of a given observation.

In order to identify influential observations whose inclusion might be problematic for the model, we look at Cook's Distance, a statistic quantifying the influence of each observation (the extent to which removing that observation would greatly modify the model fit). The formula for Cook's Distance is given in Faraway's text as: $D_i = \frac{1}{p} r_i^2 \frac{h_i}{1-h_i}$, where p is the number of regressors in the model, r_i^2 is the residual effect of observation i squared, and $\frac{h_i}{1-h_i}$ is the 'leverage term' for the observation. According to the Penn State Statistics Department webpage, "The leverage h_i is a measure of the distance between the x value for the ith data point and the mean of the x values for all n data points." Source.

- 1 #calculate/plot cook's distances to identify influential observations:
- 2 #Takes approx. 1 minute to execute
- 3 cooks_distances_plot(regress)

/usr/local/lib/python3.7/dist-packages/ipykernel_launcher.py:884: UserWarning: In Mat



NOTE: Here, the instance axis is not random, but presumably follows the progression of observations across the UTC timestamps working from March through late August, for our modeling set. This Cook's Distance plotting function's source code exists in a Python package that gets called by the function 'cooks_distances_plot()', and so whether this is true is not completely clear, but the team assumes there to have been no reason for that algorithm to have shuffled the order of observations (samples and ddms) before constructing this plot.

It appears that our model has a fair number of observations which, if removed, would greatly alter the fit of the model. The rule of thumb, according to statology.org, is that if an observation has a Cook's distance of more than $\frac{4}{n}$, where n is the number of observations, then it is a likely outlier Source. $\frac{4}{n}$ in our case is approximately equal to .0004, and so we see many such values.

Another rule of thumb, given by Cook himself, suggests that Cook's Distance values of > 1 are not of significant concern as influential values (Cited in Weisberg, Sanford. *Residuals and Influence in Regression,*. New York, Chapman and Hall, 1982.) By that rule, we would appear to have no observations whose impact on the model fit is, for that stand-alone observation, highly significant.

However, the high number of influential values under the more stringent guidelines indicates that Linear Regression may not be the best approach for modeling with this dataset and that perhaps our Machine

Learning model might give more reliable predictions.

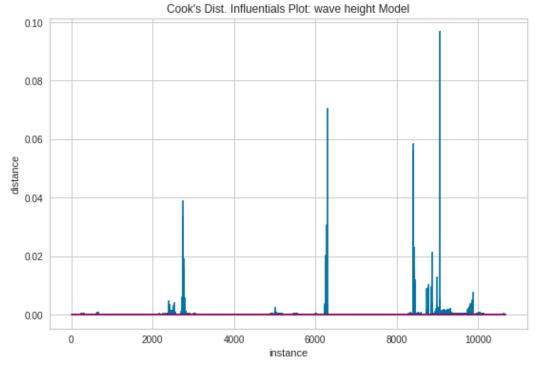
The Wave Height Model

```
1 #start by re-fitting wave height model
2 modeling frame = create dataframe(modeling set)
3 modeling frame = modeling frame.drop(['ddm average'], axis = 1)
4 modeling frame = modeling frame.drop(['wind speed'], axis = 1)
5 y = modeling frame['wave height']
6 X = modeling frame.drop(['wave height'], axis = 1)
7 \mod = sm.OLS(y,X)
8 regress = mod.fit()
1 ###conduct Bonferroni Outlier Test for our wave height Model:
2 ##NOTE: This cell requires a few minutes to run
3 bonf_test = regress.outlier_test()
1 #determine which observations might qualify as outliers; look at the number of such obs
2 bonf_outliers = bonf_outlier(bonf_test)
3 print(bonf_outliers)
         student_resid unadj_p bonf(p)
   2109
        1.053 0.292 1.0
  2117
                1.023 0.306
                                  1.0
                1.024 0.306
1.022 0.307
  2121
                                  1.0
  2125
                                  1.0
                1.037 0.300
  2129
                                  1.0
  . . .
                         . . .
  10670 1.538 0.124
                 . . .
                                   . . .
                                 1.0
  10671
                1.381 0.167
                                  1.0
                1.270 0.204
   10674
                                   1.0
  10678
                1.588 0.112
                                  1.0
                1.243 0.214 1.0
  10679
  [1563 rows x 3 columns]
1 #determine ratio of number of outliers in our model to the number of observations, base
2 observation number = 10680
3 outlier ratio = len(bonf outliers)/observation number
4 print(format(outlier ratio, '.3f'))
   0.146
```

The Wave Height model has fewer potential outliers than the wind speed model, but not by too many. There are still more of them than is comfortable for linear regression modeling, but as with wind speed, we need to see how much these translate into observations influential enough to change the model fit by their removal. We proceed to that analysis, again with Cook's Distance measurements:

1 #calculate/plot cook's distances for wave height model to identify influential observat

/usr/local/lib/python3.7/dist-packages/ipykernel_launcher.py:884: UserWarning: In Mat



Once more, using our Cook's Distances rule of thumb dictating that any distances greater than $\frac{4}{n} \approx .0004$, we see a great many influential values. However, less strict rules of thumb, some of which are also outlined at statisticshowto.com on the page 'Cook's Distance: Definition/Interpretation', dictate that only Cook's Distance values greater than .5 should be considered outliers <u>Source</u>. By that rule, we have significantly fewer for both the wind speed model and the wave height model.

In any case, we should do a little further investigation of what is happening at some of the sampling intervals where we see extreme influential value peaks.

Of particular interest is the sampling interval between roughly 8000 and 9500, as that interval of observations seems to contain the most outliers. These would correspond to DDMs from samples 2000 to 2375 in our modeling dataset, which we look at now:

```
1 outlier select = modeling set.sel(sample = slice(2000, 2375))
```

Here, it may prove informative to look at the average wind speed/wave heights for this subset of samples compared to the overall average for those variables:

```
1 compare_dependent_average(outlier_select, modeling_set)
    Average Wind Speed (Sample Subset): 7.025
    Average Wave Height (Sample Subset): 1.62
```

```
Average Wind Speed (Total Set): 6.988
Average Wave Height (Total Set): 2.066
```

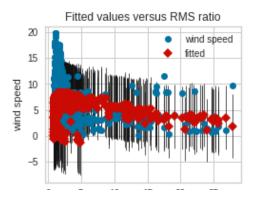
Let's see what the time intervals of the influential samples are:

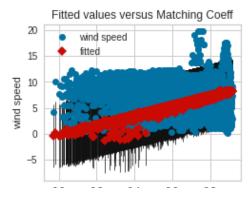
Clearly, the outlier/influential value status of so many observations in our models are linked to the much larger overall wind speed/wave height for the samples $\sim 2000-2375$. These samples correspond to the timestamps from June and July of 2021.

Analyzing Model Structure

```
1 #refit the wind speed model
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
4 modeling_frame = modeling_frame.drop(['wave height'], axis = 1)
5 y = modeling_frame['wind speed']
6 X = modeling_frame.drop(['wind speed'], axis = 1)
7 mod = sm.OLS(y,X)
8 regress = mod.fit()

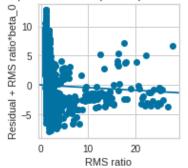
1 #fit four plots, showing the fitted values vs. the observed values for each regressor v
2 fitVsobserved(regress)
```



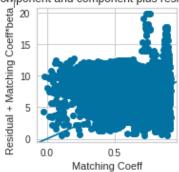


1 ##now, fit four components-plus-residuals plots, plotting residuals against each regres 2 ccpr plots(regress)

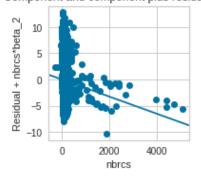




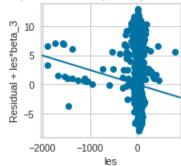
Component and component plus residual plot



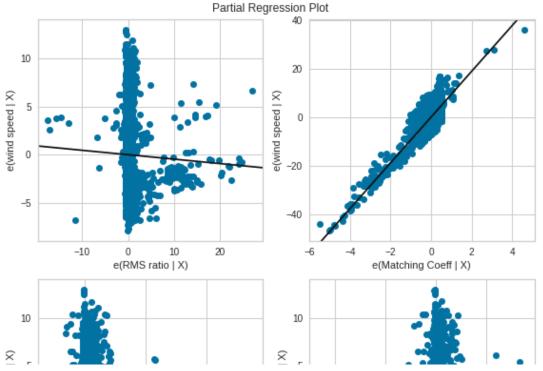
Component and component plus residual plot



Component and component plus residual plot

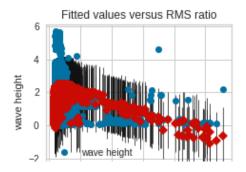


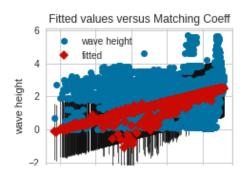
- 1 #finally, plot partial regression plots for each regressor variable:
- 2 fig = plt.figure(figsize=(8,8))
- 3 fig = sm.graphics.plot_partregress_grid(regress, fig=fig)



```
1 #do the same plotting for the wave height model
2 #first, fit the model for wave height:
3 modeling_frame = create_dataframe(modeling_set)
4 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
5 modeling_frame = modeling_frame.drop(['wind speed'], axis = 1)
6 y = modeling_frame['wave height']
7 X = modeling_frame.drop(['wave height'], axis = 1)
8 mod = sm.OLS(y,X)
9 regress = mod.fit()
```

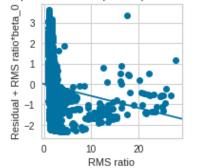
- 1 ##look at the fitted vs. observed values for each regressor variable in the wave height
- 2 #fit the wind speed model
- 3 fitVsobserved(regress)

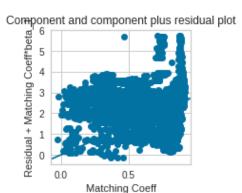




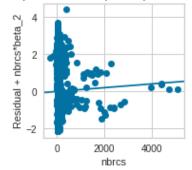
1 #look at components plus residuals plots for wave height model 2 ccpr_plots(regress)

Component and component plus residual plot

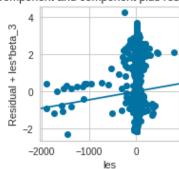




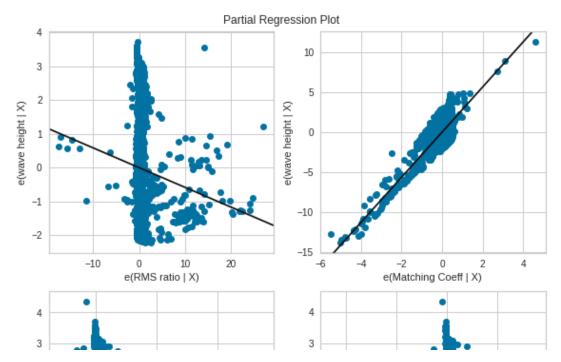
Component and component plus residual plot



Component and component plus residual plot



- 1 #finally, look at partial regression plots for each regression variable in wave height
- 2 fig = plt.figure(figsize=(8,8))
- 3 fig = sm.graphics.plot_partregress_grid(regress, fig=fig)



In both models, it appears that Maximum Template Matching Coefficient is the most significant contributor to the model fit by far. The components plus residuals plots also seem to suggest some non-linear structure in the variables, when assessed for model impact one at a time.



This, along with slight violations of error assumptions describe above, could imply that linear regression modeling is not the most beneficial technique for modeling on this dataset. Regardless, we move on to interpret the two model's fitted lines, predictions and coefficients, to wrap up the Linear Modeling segment.

Linear Model Fitting/Validation

The Wind Speed Model:

Here, we fit the final models with training/testing data for some degree of validation outside of just R^2 :

```
1 #create wind model/fit
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
4 modeling_frame = modeling_frame.drop(['wave height'], axis = 1)

1 #we use 80% of the data for training the model
2 train, test = train_test_split(modeling_frame, train_size=0.8, random_state=1)
3 modTrain = pd.DataFrame(train, columns= modeling_frame.columns)
4 modTest = pd.DataFrame(test, columns= modeling_frame.columns)

1 cols = ['RMS ratio', 'Matching Coeff', 'nbrcs', 'les']
```

```
2 \times = modTrain[cols]
3 y = modTrain['wind speed']
1 \text{ wind mod} = \text{sm.OLS}(y, x).\text{fit}()
2 wind mod.summary()
                            OLS Regression Results
      Dep. Variable: wind speed
                                       R-squared (uncentered):
                                                                0.828
         Model:
                     OLS
                                     Adj. R-squared (uncentered): 0.828
        Method:
                                              F-statistic:
                     Least Squares
                                                                1.026e+04
          Date:
                     Tue, 21 Dec 2021
                                         Prob (F-statistic):
                                                                0.00
         Time:
                     01:37:27
                                          Log-Likelihood:
                                                                -21731.
    No. Observations: 8544
                                                 AIC:
                                                                4.347e+04
                                                 BIC:
      Df Residuals: 8540
                                                                4.350e+04
        Df Model:
    Covariance Type: nonrobust
                    coef std err t P>ltl [0.025 0.975]
                  -0.0332 0.023 -1.472 0.141 -0.077 0.011
      RMS ratio
    Matching Coeff 9.3869 0.075 125.359 0.000 9.240 9.534
                  -0.0014 0.000 -4.703 0.000 -0.002 -0.001
         les
                  -0.0023 0.001 -3.441 0.001 -0.004 -0.001
       Omnibus:
                  501.961 Durbin-Watson: 1.959
    Prob(Omnibus): 0.000 Jarque-Bera (JB): 597.090
        Skew:
                 0.615
                              Prob(JB):
                                            2.21e-130
       Kurtosis: 3.405
                              Cond. No.
                                            375.
```

Warnings:

[1] Standard Errors assume that the covariance matrix of the errors is correctly specified.

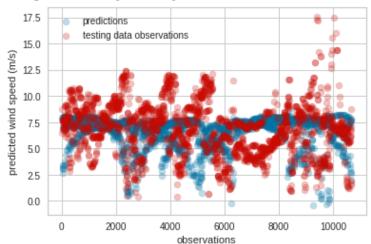
Now, we look at a simple validation of the model by calculating normalized RMSE between predicted values from trained model and dependent variable values observed in the testing data partition:

```
1 x_test = modTest[cols]
2 y_test = modTest['wind speed']
3 predictions = wind_mod.predict(x_test)
4 # Compute the root-mean-square of errors
5 rms_error = np.sqrt(mean_squared_error(y_test, predictions))
6 #normalize root mean square error
7 rms_normed = rms_error/(modeling_frame['wind speed'].mean())
8 rms_normed
0.440

1 #scatterplot of predictions values against observed values in testing data
2 fig = plt.figure()
3 ax1 = fig.add_subplot(111)
4 ax1.scatter(predictions.index, predictions, c = 'b', label = 'predictions', alpha = .25
```

```
5 ax1.scatter(y_test.index, y_test, c = 'r', label = 'testing data observations', alpha =
6 plt.xlabel('observations')
7 plt.ylabel('predicted wind speed (m/s)')
8 plt.legend(loc = 'upper left')
```





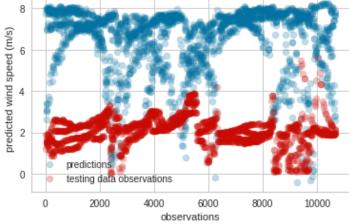
Clearly, the distribution of the predictions is far more concentrated than that of testing data observations, which falls in line with the moderate accuracy of the model as reflected by the normalized RMSE statistic.

The Wave Height Model:

```
1 ##fit final wave height model with training/testing data
2 modeling_frame = create_dataframe(modeling_set)
3 modeling_frame = modeling_frame.drop(['ddm average'], axis = 1)
4 modeling_frame = modeling_frame.drop(['wind speed'], axis = 1)
5 train, test = train_test_split(modeling_frame, train_size=0.8, random_state=1)
6 modTrain = pd.DataFrame(train, columns= modeling_frame.columns)
7 modTest = pd.DataFrame(test, columns= modeling_frame.columns)
8 cols = ['RMS ratio', 'Matching Coeff', 'nbrcs', 'les']
9 x = modTrain[cols]
10 y = modTrain['wave height']
11 wave_mod = sm.OLS(y, x).fit()
12 wave_mod.summary()
```

OLS Regression Results

```
Dep. Variable: wave height
                                    R-squared (uncentered): 0.843
                   OLS
                                  Adj. R-squared (uncentered): 0.843
        Model:
        Method:
                                          F-statistic:
                  Least Squares
                                                           1.144e+04
         Date:
                  Tue, 21 Dec 2021
                                     Prob (F-statistic):
                                                           0.00
                   01:37:28
                                       Log-Likelihood:
         Time:
                                                         -10895.
    No. Observations: 8544
                                             AIC:
                                                           2.180e+04
     Df Residuals: 8540
                                             BIC:
                                                           2.183e+04
       Df Model:
                   4
    Covariance Type: nonrobust
                                        P>ltl [0.025 0.975]
                   coef
                         std err
1 ##attempt validation with calculation of normalized RMSE between predicted values and o
2 x test = modTest[cols]
3 y test = modTest['wave height']
4 predictions = wind mod.predict(x test)
5 # Compute the root-mean-square
6 rms error = np.sqrt(mean_squared_error(y_test, predictions))
7 #normalize root mean square error
8 rms normed = rms error/(modeling frame['wave height'].mean())
9 rms_normed
   2.334
   [1] Grandard Erroro accumo shat the covariance matrix of the erroro to contoury opening.
1 #scatterplot of predictions values against observed values in testing data
2 fig = plt.figure()
3 ax1 = fig.add subplot(111)
4 ax1.scatter(predictions.index, predictions, c = 'b', label = 'predictions', alpha = .25
5 ax1.scatter(y_test.index, y_test, c = 'r', label = 'testing data observations', alpha =
6 plt.xlabel('observations')
7 plt.ylabel('predicted wind speed (m/s)')
8 plt.legend(loc = 'lower left')
   <matplotlib.legend.Legend at 0x7f86b5187a50>
```



Just as the extremely high normalized RMSE statistic calculated just above indicates, the wave height linear model is woefully inaccurate once a train/test split is used for attempted validation. This is

unsurprising, as the team assumed linear modeling on the simple calibrations we have included to be a far too simplistic a method to reliably and consistently predict something so complicated as ocean surface weather patterns.

Futher conclusions about this validation and interpretation details are offered in the section 'Conclusions' below.

```
1 del modeling_frame, regress, rms_error, rms_normed, bonf_outlier, bonf_outliers, bonf_t
2 del x_test, y_test, x, y, predictions
3 del fig, ax1
```

Machine Learning

The team spent some time experimenting with machine learning. It was recommended to the team to start with a categorical machine learning model.

The modeling dataset needs some adjustments to be ready for machine learning. Many of the variables included have no predictive power and will not be useful for machine learning. Those variables need to be removed. The timestamp variable needs to be changed into a numeric variable to be used for machine learning and last, the categories need to be generated for wind speed and wave height.

```
1 #Reach into David's file folder to extract 'ML_data.pkl'
2 ##Must reset pathTeam to eliminate Ben path info
3 pathTeam = cwd + '/drive/My Drive/'
4 #Check to add professor path
5 if os.path.exists(pathTeam + pathProfessor):
6 pathTeam += pathProfessor
7 pathTeam += David path # Should be a shortcut (Links to an external site.) to Team's sh
8 os.listdir(pathTeam)
   ['cyg firstfile sps.pkl',
    'cyg.ddmi.s20210411-010506-e20210411-171248.ll.power-brcs-full.a30.d31.nc',
    'ecmwf.t00z.pgrb.0p125.f000 2021041100.nc',
    'ecmwf.t12z.pgrb.0p125.f000_2021041112.nc',
    'ecmwf.t18z.pgrb.0p125.f000 2021031118.nc',
    'CYGNSS_0311.pkl',
    'CYGNSS 0411.pkl',
    'CYGNSS Background Collocated 20210311.nc',
    'modeling dataset.nc',
    'wValues 20210311.pkl',
    'ML data sample2.pkl',
    'wValues 20210411.pkl',
    'wValues_20210411_sample.pkl',
    'ML_data.pkl']
1 modeling set = xr.open dataset(f'{pathTeam}modeling dataset.nc')
```

```
1 ML_data_prep(modeling_set)
```

```
100% | 10680/10680 [00:00<00:00, 45870.65it/s]
100% | 10680/10680 [00:00<00:00, 64801.82it/s]
```

```
1 ML_ds = pd.read_pickle(f'{pathTeam}ML_data.pkl')
2 ML ds
```

| | ddm average | RMS ratio | Matching Coeff | wind speed | wSWH | nbrcs | les | wU10m | wV1 |
|-------|----------------|--------------|-------------------|---------------|----------|-----------|-----------|-----------|--------|
| 0 | 7401.46 | 1.077134 | 0.288958 | 9.299924 | 1.577483 | 62.245186 | 25.983896 | -8.181535 | 4.4216 |
| 1 | 14560.06 | 1.648460 | 0.756654 | 8.764927 | 1.720895 | 33.156757 | 12.924806 | -7.606535 | 4.3548 |
| 2 | 10968.92 | 1.926753 | 0.859642 | 9.552515 | 1.246971 | 38.310802 | 20.427641 | -8.074068 | 5.1048 |
| 3 | 15699.80 | 1.839177 | 0.844802 | 6.573362 | 1.902701 | 31.600922 | 13.439723 | -6.564295 | 0.3451 |
| 4 | 7512.64 | 0.963515 | 0.293512 | 8.576059 | 1.572910 | 72.594566 | -1.148427 | -7.503491 | 4.1528 |
| | | | | | | | | | |
| 10675 | 8693.80 | 1.090208 | 0.480093 | 7.691555 | 2.011177 | 60.152973 | 36.688282 | -0.149721 | 7.6900 |
| 10676 | 19592.18 | 2.520207 | 0.859517 | 8.318233 | 1.972587 | 75.276253 | 36.030926 | -3.374051 | 7.6032 |
| 10677 | 9361.48 | 1.613579 | 0.853095 | 1.934908 | 2.103230 | 88.844673 | 42.699989 | -1.882078 | 0.4490 |
| 10678 | 8453.16 | 1.104028 | 0.228713 | 6.609409 | 1.954110 | 0.000000 | 0.000000 | -1.104201 | 6.516 |
| 10679 | 8374.92 | 1.044738 | 0.347402 | 7.691287 | 2.011186 | 49.667934 | 22.902370 | -0.149276 | 7.6898 |

10680 rows x 13 columns

The new clean dataset has 10,680 observations and includes only the variables wanted for machine learning. Some small adjustments will need to be made depending on which model the team is working on (wind speed or wave height).

Wind speed model

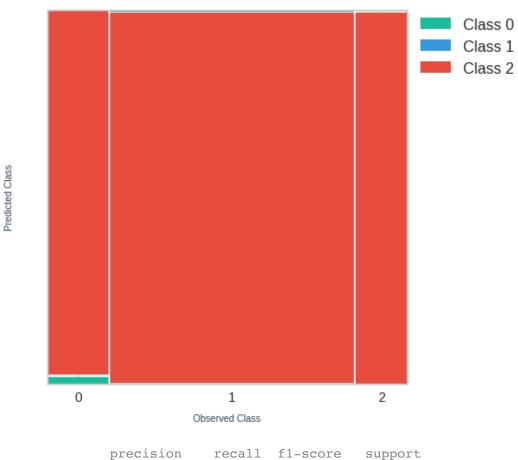
```
1 y = ML_ds['wind_category']
2
3 x = ML_ds
4 x.drop(['wU10m', 'wV10m', 'wind speed', 'wave_category', 'wind_category'], inplace=True
1 xtrain, xtest, ytrain, ytest = train_test_split(x, y, test_size = 0.15)
2
```

```
3 sgdc = SGDClassifier(max iter=10000, tol=0.01)
 4 print(sgdc)
 6 sgdc.fit(xtrain,ytrain)
    SGDClassifier(max iter=10000, tol=0.01)
    SGDClassifier(max iter=10000, tol=0.01)
 1 score = sgdc.score(xtrain, ytrain)
 2 print(f"Training score: {score}")
4 null = max(ytest.value_counts())/sum(ytest.value_counts())
 5 print(f"Null training score: {null}")
 6 print()
7
8 ypred = sgdc.predict(xtest)
10 cm = confusion matrix(ytest, ypred)
11 CM=pd.DataFrame.from dict({
      'Calm actual': [cm[0][0], cm[0][1], cm[0][2]],
13
      'Mild_actual': [cm[1][0], cm[1][1], cm[1][2]],
      'Strong actual': [cm[2][0], cm[2][1], cm[2][2]]
14
15 },
16 orient='index', columns=['Calm predict', 'Mild predict', 'Strong predict'])
17 print(CM)
18 print()
19
20 results = [
      [cm[0][0], cm[0][1], cm[0][2]], # predictions for class 1
21
      [cm[1][0], cm[1][1], cm[1][2]], # predictions for class 2
22
23
      [cm[2][0], cm[2][1], cm[2][2]], # predictions for class 3
24 ]
25
26 nclass_classification_mosaic_plot(3, results)
27 print()
28
29 cr = classification report(ytest, ypred)
30 print(cr)
```

Training score: 0.1440846001321877 Null training score: 0.6841448189762797

| | Calm_predict | Mild_predict | Strong_predict |
|---------------|--------------|--------------|----------------|
| Calm_actual | 5 | 0 | 270 |
| Mild_actual | 4 | 0 | 1092 |
| Strong_actual | 0 | 0 | 231 |

Mosaic Plot of Confusion Matrix



precision recall f1-score

Significant Wave Height Model

```
1 ML ds2 = pd.read pickle(f'{pathTeam}ML data.pkl')
3 y2 = ML ds2['wave category']
5 x2 = ML ds2
6 x2.drop(['wU10m', 'wV10m', 'wSWH', 'wave_category', 'wind_category'], inplace=True, axi
     _warn_pri(average, mourrier, msy_start, ren(resurt))
1 xtrain2, xtest2, ytrain2, ytest2 = train_test_split(x2, y2, test_size = 0.15)
3 sgdc2 = SGDClassifier(max iter=10000, tol=0.01)
4 print(sgdc2)
6 sgdc2.fit(xtrain2,ytrain2)
```

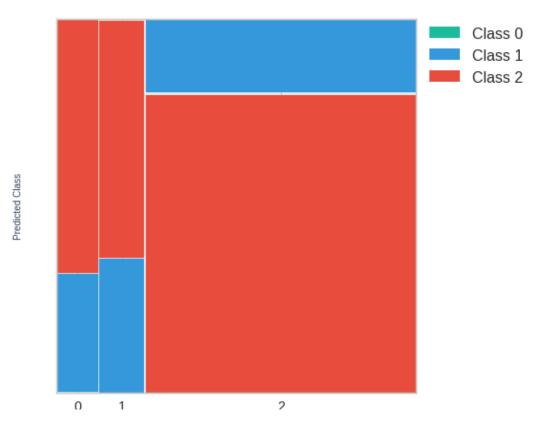
```
SGDClassifier(max_iter=10000, tol=0.01)
SGDClassifier(max_iter=10000, tol=0.01)
```

```
1 score = sgdc2.score(xtrain2, ytrain2)
 2 print(f"Training score: {score}")
 4 null = max(ytest2.value_counts())/sum(ytest2.value_counts())
 5 print(f"Null training score: {null}")
 6 print()
 7
 8 ypred2 = sgdc2.predict(xtest2)
10 cm2 = confusion matrix(ytest2, ypred2)
11 CM2=pd.DataFrame.from dict({
      'High_actual': [cm2[0][0], cm2[0][1], cm2[0][2]],
13
       'Low_actual': [cm2[1][0], cm2[1][1], cm2[1][2]],
14
      'Medium actual': [cm2[2][0], cm2[2][1], cm2[2][2]]
15 },
16 orient='index', columns=['High_predict', 'Low_predict', 'Medium_predict'])
17 print(CM2)
18 print()
19
20 results = [
      [cm2[0][0], cm2[0][1], cm2[0][2]], # predictions for class 1
22
      [cm2[1][0], cm2[1][1], cm2[1][2]], # predictions for class 2
      [cm2[2][0], cm2[2][1], cm2[2][2]], # predictions for class 3
23
24 ]
26 nclass_classification_mosaic_plot(3, results)
27 print()
28
29 cr2 = classification report(ytest2, ypred2)
30 print(cr2)
```

Training score: 0.6560916501432034 Null training score: 0.7590511860174781

| | High_predict | Low_predict | Medium_predict |
|---------------|--------------|-------------|----------------|
| High_actual | 0 | 59 | 127 |
| Low_actual | 0 | 72 | 128 |
| Medium_actual | 0 | 239 | 977 |

Mosaic Plot of Confusion Matrix



The quality of the model can be determined by looking at the 4 different parts of the printout. The first is a training score for the model and a null model, the sceond is a confusion matrix, the third is a mosaic plot of the confusion matrix, and the final part is a classification report.

The training score measures the number of correct guesses against the total number of guesses. This number is not very useful if not compared to a null score. The null score is the training score a model would get if it only guessed the most common class. In general, it is best for the model training score to be greater than the null training score, however, this is not always true. Due to that, the training score is not the best method for determining the quality of a model.

The next part of the printout is a confusion matrix. A confusion matrix tallies what the correct answer of a given test element is vs what the guess from the model was. If the model guesses correctly, a tally will be made along the main diagonal. For example, if the test element belongs to category 1 and the model guesses category 1, the count in position (1,1) on the matrix will go up by one. If instead, the maching guesses category 2, the matrix tally will increase by one in the position (1,2). For a confusion matrix to provide evidence that demonstrates a good model, the elements of the matrix not on the main diagonal should be very low as they represent incorrect guesses by the model.

The mosaic plot of the confusion matrix is a way to visualize the confusion matrix. The width's of the classes on the x-axis are determined by the proportion of each class in the testing dataset. The bars are colored based on how many times each class was guessed. For a model that is accurately classifying observations, one would expect the bars to be colored mostly by the proper prediction class color (i.e. the bar of observed class 0 would be mostly the color for class 0 as defined in the legend).

The final part of the printout is a classification report which gives an in-depth look into how a classification model performed. The 'precision' column gives the ratio of correct guesses for an individual class to total guesses of that class. The 'recall' column gives the ratio of correct guesses of a class to total occurances of that class. The 'f1-score' column is a metric that combines precision and recall and is used to compare models, not to determine model accuracy. The f1-score is given by $f1 = \frac{2*Recall*Percision}{(Recall+Precision)}.$ The final column 'support' counts the total number of times that class appears in the testing data.

Depeding on how the dataset gets split into training and testing data, the results will vary. For every trial the team ran, the results were not encouraging.

In general, the windspeed model was not able to beat the null training score. However, there was one trial the team ran that beat the null training score. The confusion matrix consistantly shows too many incorrect guesses. The mosaic plot shows that the class 'mild' was guessed almost every time. The classification report shows a poor precision and recall score. All of these factors suggest that the windspeed model created is not a good model. The column with the most support (Mild) is consistantly the most accurately guessed. This leads the team to believe that increasing the size of the modeling dataset would likely improve the model.

In general, the wave height model showed better restults as the training score was very close to the null training score and depending on how the dataset is split for training and testing, the training score was actually higher than the null score on occasion. That being said, the confusion matrix still showed a signifant number of incorrect guesses and the mosaic plot shows that the 'medium' category was almost exclusively guessed. The classification report backed up the findings of the confusion matrix, suggesting the model is not accurately classifying the different types of significant wave height. The team believes that increasing the size of the modeling dataset would also improve this model.

Conclusions

• What are you taking away from your work?

Linear Modeling Conclusions:

The error assumptions of linear modeling were only partly satisfied by our dataset.

Our most interesting conclusions for the linear model came where we assessed the status of influential values/outliers, and also where we performed model fitting/prediction. With regard to outliers/influential

values, the consecutive times with the most influential values might suggest that there could be a seasonal component to wind speed/wave height, and that either a linear model with a much larger range of dates contributing samples, or a model accounting for seasonality, could reduce the status of these observations as influential values. This could be a promising avenue of future research.

Linear Modeling Results/Interpretation:

The Linear Model we fit for Wind Speed gives the fitted equation:

$$windspeed = -.0332(RMSratio) + 9.3869(MatchingCoeff) - .0014(nbrcs) - .0023(les)$$
 The Linear Model we fit for Wave Height gives the fitted equation:

$$wave height = -0.0539 (RMSratio) + 2.8247 (MatchingCoeff) + .0000771 (nbrcs) + 0.0003 (logical content for the content for t$$

For each respective model, assuming we can count on the assumptions of linear modeling for this dataset, there is a 95% probability that the confidence intervals listed in the regression summaries above contain the true values of the β coefficients.

According to Statology.org, the generally closer a Normalized RMSE value is to zero, the more reliable the model is (the less likely to be overfit or underfit). Several university-affiliated postings on ResearchGate.net suggest that an NRMSE of >= .5 reflects a generally inability of the model to predict reliably Link By that standard, our Wind Speed Linear Regression Model predicts with greater reliability than our Wave Height Model.

Machine Learning Modeling Conclusions:

The team was not able to develope a categorical machine learning model capable of predicting wind speed or significant wave height. While neither model was successful, the significant wave height model generally has an accuracy score around 80%. It is possible that the model could be improved to the point that the accuracy score reaches a desirable level.

What do you want the reader to take away?

Hopefully, the reader/notebook user has taken away not only more insight into the CYGNSS project by NASA, but has been stimulated to consider future directions modeling research with that CYGNSS data may take. We especially hope this notebook has contributed an interesting combination of DDM calibrations to predict weather patterns on the ocean surface in various simple models. It remains for further research to enhance the sophistication of the CYGNSS modeling work begun here, and to build up the accuracy/reliability of predictions.

Be honest about what conclusions are really supported

Linear Modeling Limitations:

While the ability to use linear regression modeling to predict wind speed/wave height from Maximum Template Matching Coefficient of DDM data seemed promising, the model did, realistically, have a worrying degree of violation of linear modeling assumptions. Thus, the interpretation of OLS coefficient estimates from our model, as given above, should be taken with a large grain of salt.

Machine Learning Modeling Limitations:

There were many limitations with the teams' efforts in developing a categorical machine learning model. The team was limited by its understanding of the SGDClassifier. More optimal parameter tuning for this model may exist. The team attempted to find the optimal parameter settings by trial and error. Therer is also a significant amount of data available in the collocated data base to add to the training set if a future team is to try to improve the models. There could also be room for improvment if more time were to be spend on the variable selection process. Taking any, or all of these steps would likely improve the categorical machine learning model.